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# Explosive sequence of La Soufriere St Vincent April 2021: insights into drivers and consequences via eruptive products

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1 **Explosive sequence of La Soufrière St Vincent April 2021: insights into drivers and consequences**  
2 **via eruptive products**

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14

15 **Abstract**

16 This paper forensically reconstructs the timings, impacts and processes that drove the sequence of  
17 explosive eruptions of La Soufrière, St Vincent in April 2021 using a combination of field-based  
18 stratigraphy and textural dissection of the deposit character together with contemporary visual  
19 observations.

20 Explosive activity on 9<sup>th</sup> and early on 10<sup>th</sup> April involved destruction of almost all of the 2020/2021  
21 lava dome, ~ 60% of the 1979 dome and formation of a 600 m diameter crater by 2pm UTC on 10<sup>th</sup>  
22 April. Following the initial explosion, plumes rose to altitudes of ~15 km and pyroclastic density  
23 currents (PDCs) formed by column collapse, first occurred on 10<sup>th</sup> April, only after > 24hrs of  
24 explosive activity. Dense PDCs reached the sea only in the Larikai and Roseau Valleys, and dilute  
25 PDCs were restricted to within 2.5 km of the Summit Crater rim.

26 The tephra fallout deposits are stratified, composed of numerous layers of both lapilli-rich and ash-  
27 rich layers, which we have grouped into at least 7 Units, based on their common characteristics  
28 (Units 1 to 7).

29 Volume estimates, using a range of techniques to constrain uncertainties, indicate that the bulk  
30 volume of tephra (fallout and PDC) is  $1.19 \times 10^8 \text{m}^3$  +/- 20% making this a VEI 4 eruption.

31

32

33

34 **1. Introduction**

35 The island of St Vincent lies in the southern part of the Eastern Caribbean volcanic arc, and La  
36 Soufrière volcano sits at the northern end of the island (Fig 1). The volcano rises to 1204 m and 1.5  
37 km diameter crater cuts the summit, hereafter referred to as the Summit Crater. It is the most  
38 active terrestrial volcano in the region. Having experienced four historical explosive eruptions, in  
39 1718, 1812, 1902 and 1979, and at least one solely effusive eruption in 1971 (Robertson 2005).  
40 Radiocarbon dating has established that there were at least two prehistoric eruptions in the last  
41 1000 years (prehistoric time is < 1700 CE in the Eastern Caribbean), one ~1580 CE, and another  
42 ~1440 CE (Cole et al. 2019).

43 All the explosive eruptions at La Soufrière in the last 1000 years have generated pyroclastic density  
44 currents (PDCs), however these have varied considerably in their extent. The 1902 eruption  
45 generated extensive, broadly radially distributed PDCs, resulting in the deaths of ~1500 people in  
46 settlements around the volcano (Anderson and Flett 1903). Also, the 1812 eruption generated PDCs  
47 in several valleys, particularly to the southwest, leading to fatalities (Smith, 2011). The 1979 eruption  
48 formed much smaller PDCs which, apart from in proximal regions, were confined to valleys to the  
49 west and southwest, draining the lowest part of the Summit Crater rim (Shepherd et al. 1979).  
50 Tephra fallout was extensive with all explosive eruptions, with at least the 1812, 1902 and 1979  
51 eruptions generating ashfall in Barbados 170 km to the east (Anderson and Flett 1903, Shepherd et  
52 al. 1979). The prehistoric events in 1580 and 1440 CE were apparently larger and more intense,  
53 generating significant lapilli fallout deposits on the island prior to the onset of PDC activity (Cole et al  
54 2019).

55

### 56 **The April 2021 Explosive activity**

57 Eruptive activity at La Soufrière, St Vincent restarted after a > 40 year hiatus on 27<sup>th</sup> December 2020  
58 following a minimal period of precursory seismicity (Joseph et al. 2022, Thompson et al. 2022).  
59 Passive extrusion of a basaltic andesite lava dome then took place for ~ 3 months in the  
60 southwestern part of the Summit Crater (Stinton et al. this volume). Seismicity changed on 23 March  
61 2021 with a short VT (volcano tectonic) earthquake swarm (226 events), a second more intense VT  
62 swarm occurred on 5<sup>th</sup> April, (476 events) followed by banded tremor on 8<sup>th</sup> April. Explosive activity  
63 started on the 9<sup>th</sup> April 2021 (Joseph et al. 2022), although this commenced with a single explosion  
64 at 12:41 UTC, the rapidity with which explosions occurred (~ hourly), together with their semi-  
65 continuous nature through the evening of the 9<sup>th</sup> April and for the following 48 hours made  
66 immediate interpretation of the nature of these explosions difficult. Much of the northern part of  
67 the island remained in darkness on 10 and 11<sup>th</sup> April, owing to the presence of extensive, semi

68 continuous ash plumes. Following this, explosions continued at a lower frequency until 22<sup>nd</sup> April  
69 (Joseph et al. 2022).

70 This paper uses a combination of field-based stratigraphy of the tephra sequence, contemporary and  
71 visual observations, along with more detailed analysis of the fallout and PDC deposits themselves to  
72 reconstruct the changing nature of the explosions and their relative timing. We offer some  
73 interpretations of the critical changes that influenced explosion style and impact. The interplay  
74 between the initial overburden of the dome, the likely ingress of gassier, more buoyant magma into  
75 the conduit system, and the interactions between the vent region, accumulating Summit Crater fill  
76 deposits and the prevailing meteorological conditions all influenced the differing character of these  
77 explosions.

78

## 79 **2. DEPOSIT CHARACTER**

### 80 **The Tephra sequence**

81 Documentation of the tephra deposits was made during three field seasons. Distal products > 4 km  
82 from the Summit Crater were documented immediately following cessation of the explosive activity,  
83 between 25<sup>th</sup> April and 10<sup>th</sup> May 2021. Timed tephra sampling of measured areas was made at a one  
84 location on St Vincent and two on Barbados during the explosive activity on 10 - 13<sup>th</sup> April. Two  
85 further field campaigns in early 2022 documented more proximal products between 4 km and the  
86 Summit Crater rim. Overall the tephra deposits were investigated at more than 80 locations around  
87 the island. In general, the sequence is stratified, composed of numerous layers, defined by variations  
88 in the grainsize and nature of the components. The sequence reaches >70 cm in thickness 500m  
89 southeast, and > 1 m ~ 800 m southwest of the Summit Crater. Several 'Units' were identified, from  
90 1 to 7 from the base upwards, each Unit is composed of a number of layers with similar  
91 characteristics that we were able to recognise and correlate around the volcano (Fig 2).

92 Apart from in the Larikai and Roseau valleys (on the West flank) and in the proximal regions < 2.5 km  
93 from the Summit Crater rim, where PDCs occurred, deposits comprising all Units described below  
94 are considered to be fallout. PDCs were generated contemporaneously with some of these Units,  
95 and these products are described in a later section.

96 Below we describe the different features of the various fallout Units. Key observations are  
97 summarised in Table 1.

98

**100 Unit 1 (U1)**

101 The lowermost unit, U1, is a moderate grainsize lapilli deposit with a maximum thickness of 20 cm,  
102 350 m southeast of the Summit Crater rim (Fig 2 and 3). At many localities it shows reverse grading,  
103 defined by the lowermost third of the unit being slightly finer grained than the upper part. This is  
104 particularly evident on the eastern side of the volcano, where the Unit is preferentially dispersed. . A  
105 crude stratification, defined by small variations in grainsize of the lapilli was present, such that at  
106 least three sub-units could be identified (Fig 3 and 3a). Local horizons of hydrothermally altered  
107 fragments (up to ~20 wt%), are found close to the change in grainsize between the lower (finer-  
108 grained) and the upper (slightly coarser) parts of the unit. The contact with the overlying Unit 2 is  
109 typically defined by a fine grained laminated brown ash, a few mm thick, possibly representing a  
110 pause in deposition between Units.

111

**112 Unit 2 (U2)**

113 This Unit is a notably ash-rich, stratified deposit that reaches a maximum of 22 cm thick, 800 m east  
114 of the Summit Crater rim. The stratification is defined by the presence of numerous thin lapilli  
115 layers. In distal regions these were sometimes only a few mm thick (Fig 2 and 3). U2 shows a distinct  
116 colour change from brown (lower part) to grey (upper part).

117 Although generally ash-rich, at proximal locations the brown lower part of U2 is finer grained,  
118 whereas the upper part contains more abundant and coarser lapilli layers that are notably coarser  
119 than those lower in this Unit (Fig 2 and 3a). At distal locations > 4 km from the Summit Crater, these  
120 layers are typically horizons of lapilli only a clast thick, typically < 1 cm in diameter, however, notably  
121 larger lapilli of clasts 1 - 3 cm in size, both of pale vesicular scoria and dense lithologies are locally  
122 observed throughout the Unit. Up to seven individual layers were identified within the Unit, which  
123 are particularly well-defined at upwind locations (Fig 2 and 3b).

124

**125 Unit 3 (U3)**

126 This Unit is a distinctive coarse-grained, double lapilli layer couplet (L1 and L2) up to 32 cm thick, 200  
127 m from the south-eastern Summit Crater rim, although in localities > 4 km it is typically ~2 cm thick.  
128 It forms a key marker horizon within the tephra sequence around much of the island (Fig 2 and 3).

129 In proximal areas, < 1km from the Summit Crater rim, the two lapilli layers are quite distinct (Fig 2  
130 and 3c). The lowermost lapilli (L1), was up to 12 cm thick and typically contained a thin fine ash  
131 layer up to 1 cm thick, in the upper two thirds of the lapilli. Crude reverse grading is evident across  
132 the whole layer. The uppermost lapilli (L2) is a normally graded lapilli layer up to 8 cm thick. An ash  
133 layer, up to 7 cm thick, rich in accretionary lapilli separates the two lapilli layers, however in distal  
134 regions (> 4 km from the Summit Crater) this ash layer is typically < 0.5 cm thick.

135 This Unit forms one of the coarsest lapilli layers of the sequence with outsized lapilli up to 5 cm in  
136 diameter in distal regions that sometimes protruded through the overlying Unit 4 and into Unit 5.  
137 (Fig 2 and 3)

138

#### 139 **Unit 4 (U4)**

140 This Unit is a single ash layer up to 12 cm thick in the most proximal locations. At more distal  
141 localities > 4 km from the Summit Crater rim this Unit is just a few mm thick. Small aggregates,  
142 including accretionary lapilli, up to 8 mm diameter are abundant in this unit (Figs 2 and 3d).  
143 Scattered coarser lapilli of both pale scoria and dense lithologies, < 5 mm diameter are also present  
144 in this Unit.

145

#### 146 **Unit 5 (U5)**

147 The uppermost distinctive lapilli layer of the tephra sequence, this unit formed the surface carapace  
148 lapilli on the eastern (windward) side of the volcano. It is a maximum of 5 cm thick in proximal  
149 locations to the southeast, although thicknesses were difficult to estimate on the Eastern flank  
150 owing to it being the uppermost, uneven lapilli on the tephra deposit surface. Vesicular scoria lapilli  
151 and blocks are typical however, dense glassy non-vesicular clasts are also notable, giving the coarse  
152 lapilli an apparently bimodal vesicularity. In the southwestern region U5 is up to 2 cm thick and was  
153 overlain by the fine ash of Unit 6 (Fig 2 and 3)

154

#### 155 **Unit 6 (U6)**

156 This unit is found solely on the south-western side (Leeward side) of the volcano and is formed by a  
157 number of ash layers. Collectively these form thicknesses of up to 20 cm, 2 km to the southwest of  
158 the Summit Crater, but individually these are up to 10 cm thick (Fig 1 and 2). These ash layers are

159 notably fine grained and rich in accretionary lapilli. Individual accretionary lapilli are typically < 1 cm  
160 in size.

161 It is evident that some of the thickest and most abundant of these ash layers is close to the Roseau  
162 valley and upper Wallibou valleys down which PDCs travelled.

163

#### 164 **Unit 7 (U7)**

165 Similar to U6, this Unit is found only in a few localities on the south-western side of the volcano,  
166 such as around the Dry Wallibou valley. It is formed by 3 or 4 thin, relatively fine-grained lapilli layers  
167 that are collectively up to 4 cm thick. Individual layers are 0.5-1.5 cm in thickness (Fig 2).

168

#### 169 **Pyroclastic Density currents**

170 Pyroclastic density currents (PDCs) were associated with the explosive activity and entered several  
171 valleys draining the Summit Crater. They reached the sea in the Larikai and the Roseau valleys on the  
172 western and south-western flank of the volcano only (Fig 4a).

173 PDCs also extended down several of the valleys leading from the southern part of the Summit Crater  
174 rim and into the upper Wallibou valley, reaching as far as Trinity falls / Wallibou hot springs,  
175 approximately 2.5 km from the Summit Crater rim. PDCs also extended short distances  
176 (approximately 1 km) down valleys to the southeast, effectively the head of the Rabacca valley.

177 In proximal regions on ridges and valley sides, up to 2.5 km from the Summit Crater rim, vegetation  
178 was removed and trees were extensively felled in directions typically away from the crater(Fig 4a,5a  
179 and b). . Thus, we infer that dilute, low particle concentration PDCs were responsible for this  
180 destruction and have used the presence of felled trees to map their extent (Fig 4a). Dissected  
181 deposits of the valley-filling PDCs in the Roseau and Larikai valleys, and in proximal areas on the  
182 south-eastern flank were documented (Fig 4 b-e).

183

#### 184 **Larikai and Roseau valley PDC deposits**

185 At the mouth of the Larikai Valley PDCs form a fan of deposits, 300 m wide and up to 20 m metres  
186 thick (See Fig 4b). Wood contained within the deposits is carbonised and local gas escape structures  
187 confirm a primary PDC origin for these deposits. A series of at least six massive PDC units form the  
188 Larikai fan. The lowermost unit contains regions locally enriched in coarse clasts, composed of

189 moderately vesicular and dense glassy blocks of lava up to 3 m in diameter. Dissection of these  
190 deposits show these concentrations define convex lobate surfaces within the units. The lowermost  
191 deposits form a series of four units, without ashfall layers between them. Overlying this, three  
192 massive PDC units, interbedded with an accretionary lapilli-rich fallout layer up to 11 cm thick and a  
193 series of stratified accretionary lapilli bearing layers, collectively up to 90 cm thick (Fig 4b). A  
194 massive poorly sorted layer ~ 1 m thick, overlies the accretionary lapilli-rich fallout layer, and a  
195 stratified sequence of accretionary lapilli bearing layers collectively up to 90 cm thick occurs higher  
196 in the sequence ('S' on Fig 4b). When traced laterally these layers show notable lateral thickness  
197 variations and collectively this sequence forms large wavelength dune-bedforms. These are  
198 therefore considered to represent alternating thin PDC and fallout layers. A massive unit, up to 2 m  
199 thick, caps the sequence. Approximately 1 km up the valley the total thickness of PDC deposits  
200 reach thicknesses of up to 30 m.

201

202 Similar to the Larikai valley, dissection of deposits reveals that the Roseau valley contained several  
203 massive flow units (Fig 4c). Up to four separate units, ranging from 2 to 8 m thick individually, are  
204 present reaching a maximum of 20 m in thickness, although more typically around 12 m. All these  
205 units contain both cauliflower vesicular and dense glassy scoria clasts up to ~50 cm in diameter. A thin  
206 15 cm thick accretionary lapilli bearing ash fall layer is interbedded between the flow units ('A' on Fig  
207 4c) which we correlate with the Unit 6 of the tephra fallout.

208 Tree branches, trunks and other vegetation contained within these deposits was generally fully  
209 carbonised, although locally some wood debris had remained uncharred.

210

### 211 **Upper Southeastern flank deposits**

212 On the proximal south-eastern flanks of the volcano in the region near Jacobs Well (~0.5 - 1 km from  
213 the Summit Crater rim) a series of massive, poorly sorted, valley-filling deposits up to at least 6 m  
214 thick, are observed partially filling valley depressions. Critically these massive deposits are  
215 interbedded between several tephra fallout layers, identified as Units 3, 4 and 5 (See Fig 4d). Fallout  
216 of Units 1 and 2 are present at the base of the sequence below the PDC deposits.

217 Vegetation and wood fragments incorporated within these massive valley-filling deposits were  
218 abraded and debarked but generally not carbonised. Nevertheless, these deposits are considered to  
219 be primary PDC deposits.



220

### 221 **Dilute PDC deposits**

222 Deposits formed by the dilute PDCs that caused extensive tree felling above the valley thalwegs (on  
223 slopes and ridges) were notably variable. Within 1 km of the southeast of the Summit Crater rim  
224 deposits were, < 15 cm thick, composed of fine lapilli, and typically poor in fine ash. These were  
225 identified below the U5 lapilli layer and these deposits were relatively continuous, but showed some  
226 minor local lateral thickness variations.

227 To the southwest of the Summit Crater rim the deposits on ridges were thicker and more extensive.  
228 A series of deposits interbedded with U3, 4 and 5 are individually 10 – 25 cm thick and up to 40 cm  
229 thick < 1 km from the Summit Crater rim and show extensive thin < 5 cm thick lenses of ash poor,  
230 fine grained lapilli within more ash rich deposits (Fig 5c). Other units are formed by variable  
231 thickness ash poor, well-sorted fine lapilli layers. Locally these deposits show crude dune bedforms  
232 and cross stratification with marked thickness variations.

233 The numerous trees felled by these dilute PDCs show strong abrasion, typically only on the volcano  
234 facing sides. Locally trees with trunks > 50 cm in diameter often remained standing (Fig 5 a and b).  
235 Many trees were bent over or snapped at heights of 2-3 m above the base (Fig 5a). Carbonisation of  
236 vegetation inundated by passage of these dilute PDCs was absent, although tree fern trunks  
237 exhibited a slight blackening of their exterior. Similar lack of carbonisation of trees impacted by PDCs  
238 has been reported associated with the 18<sup>th</sup> May 1980 Mt St Helens blast PDC (Moore and Sission  
239 1981), the 2008 eruption of Chaiten (Swanson et al 2013, Major et al 2013) Chile, and at Kelud in  
240 Java related to the 2014 eruption (Maeno et al 2019). At Mt St Helens and Chaiten temperatures for  
241 uncharred trees were inferred to result from shortlived PDCs < 200°C.

242

## 243 **3. CHANGING DEPOSIT CHARACTERISTICS**

### 244 **Grainsize analyses**

245 Particle size analyses were undertaken on ~ 60 samples in order to characterize the grainsize of  
246 tephra and its variation through the explosion sequence. Mass fractions were acquired through  
247 manual sieving all of clasts down 125 µm, and then using a Retsch Particle Size and Shape Analyzer  
248 CAMSIZER X2 for a sub-sample of all material < 125 µm. We focussed on obtaining samples of the  
249 different Units through tephra sequences at a range of locations (Fig 6).

250 Cumulative grainsize distributions show a clear transition between U1 and 2, with U1 fallout  
251 deposits being notably fines-poor and U2 deposits showing a distinctly more ash-rich character (Fig 6  
252 a). Unit 3 deposits (green lines on Fig 6a) have a varied grainsize, generated either by coarser lapilli  
253 or fine ash layers. U5 plots on the coarse side of lapilli.

254 The field-based subdivision of deposits into either lapilli or ash hold true in the cumulative grainsize  
255 distributions plots (Fig 6b) albeit with some overlap. Similarly, deposits identified as dilute PDCs  
256 (yellow lines on Fig 6 b) are better sorted and poorer in fine ash than other deposits.

257

258 Grainsize data through the tephra sequence show a slight coarsening upwards vertically through U1,  
259 (Sandy Bay school Md and d) - Fig 6 c) reflecting the visible reverse grading, with continued  
260 coarsening through U2. Lapilli layers of U3 or U5 form the coarsest fallout deposits of the sequence.  
261 This is corroborated by maximum clast size data (five largest clasts of both vesicular scoria and dense  
262 clasts, Figure 6c and d) . There is a general increase in maximum clast size through the sequence,  
263 with the coarsest clasts (particularly scoria) occurring either in U3 or U5.

264 Ash aggregates, including accretionary lapilli, were abundant from U3 onwards. Accretionary lapilli  
265 were particularly abundant on the leeward (west) side of the volcano. The presence of accretionary  
266 lapilli has impacted the grainsize distribution of deposits (Fig 6 c), and is particularly notable in ash-  
267 rich deposits on the Leeward trail where peaks at 1 phi fraction in U4 and U6 arise from ash  
268 aggregation (Fig 6c).

269

## 270 **Component Analyses**

271 The 2021 tephra deposits contain a wide range of components which we initially classified in the  
272 field as pale scoria (40-60% vesicles), semi vesicular (denser) scoria (20-40% vesicles), dense  
273 (sometimes glassy) clasts (<20% vesicles), hydrothermally altered fragments and free crystals. This  
274 classification was also used to broadly classify sampled fragments >1mm using the binocular  
275 microscope (Fig 7a-d).

276 Pale grey vesicular scoria, representing the juvenile magma, occurs in all units of the tephra  
277 sequence in varying quantities. The lowermost parts of U1 contain only 10% pale scoria (>1mm) but  
278 this increases to ~ 40 wt% of clasts in the upper part of U1. Within U2 this remains at ~ 40 wt% and  
279 in U3, lapilli layers contain up to 60 wt % pale vesicular scoria.

280 Dark coloured, semi- vesicular and dense clasts, (Fig 7 b and c) are most abundant in U1, forming up  
281 to 40 and 56 wt % of components > 1mm respectively. Dense clasts were present in all units forming  
282 a significant proportion (up to 40%) of U3 and U5. Hydrothermally altered clasts occurred mainly  
283 within U1 (Fig 7d). Some horizons of U1 contained up to 20 wt % hydrothermally altered fragments.

284 We also determined the density of sampled vesicular clasts (both pale and denser scoria) in the 16-  
285 32 mm range using the Archimedes principle to calculate the bulk density of these clasts (Fig 8).  
286 These demonstrate that there are significant variations in both mode and range of vesicularities  
287 between the Units. Clasts distinguished in the field or by binocular microscope form a continuum but  
288 generally samples classified as denser scoria have vesicularities < 30%. U1 contains a broad range of  
289 vesicularities with a mode at 25-30% vesicles. The vesicularity increases markedly to 40-45% in U2  
290 and 45-50% vesicles in U3 and U5, which have similar vesicularity ranges.

291 Vesicle textures were further interrogated via quantitative analysis of BSE images for scoriaceous  
292 and semi vesicular clasts from U1, U2 and U5, and compared qualitatively with images of two clasts  
293 (semi-vesicular and scoriaceous) from U3. BSE images were acquired on a Zeiss Gemini 300 SEM at  
294 the University of East Anglia, a nested imaging strategy was applied after Shea et al., (2010) to  
295 capture the full range of vesicle sizes (see appendix for image locations and sizes). Key  
296 measurements including minimum pixel size are shown in Table 2. Vesicle shape and size was  
297 quantified using *Circularity* and *Roundness* parameters as defined by Liu *et al.*, (2015, the aspect  
298 ratio of the best fit ellipse and vesicle area (plotted in Fig 9 a - f). These parameters highlight key  
299 differences between U1/U2 and U5 and agree with qualitative observations of BSE images. Density  
300 plots (the representation of the relative occurrence of measurements) of vesicle area against aspect  
301 ratio (Fig. 9a – c) show that U1 and U2 have two modes in their vesicle area populations and extend  
302 to (a) smaller areas and (b) higher aspect ratios than U5, though most vesicles across all units are  
303 relatively unsheared. Density plots of *Roundness* vs *Circularity* (Fig 9d – f) show more subtle  
304 differences between the bubble populations. In U5, *Circularity* and *Roundness* plot much closer to  
305 each other than U1 and U2. This indicates that for the U5 clasts analyses surface roughness exerts  
306 less of a control on bubble shape than elongation, in contrast to U1 and U2 where the increased role  
307 for surface roughness is consistent with a higher pre-existing proportion of microlites on  
308 vesiculation. BSE images of bubble textures in U1, U2 and U5 are shown below shape plots (Fig.9 g –  
309 i). Images from a semi-vesicular (Fig.9j) and a vesicular (Fig. 9.k) clast from U3 show vesicle textures  
310 intermediate between U2 and U5, with the scoriaceous clast closer to U5 and the semi-vesicular  
311 clast more closely resembling U2. A similar pattern is seen in the microlite populations of the  
312 analysed clasts (Frey *et al.*, this volume). Bubble number densities (BNDs) (corrected to remove  
313 phenocryst phases) in 2D were calculated using the analysed images, the BND of U5 is approximately

314 half that of U1 and U2, despite having the highest vesicularity (39% in 2D). This confirms the  
315 qualitative observation that U5 has a notably larger vesicle population than either U1 or U2, but a  
316 more detailed analysis is provided in Christopher et al., (this volume).

317

### 318 **Volume of the explosive phase**

319 In order to estimate the volume of the products of the explosive activity three different components  
320 were considered: a) tephra fallout deposits from isopach maps b) PDC deposits in valleys outside the  
321 Summit Crater and c) the intracrater fill.

322

### 323 **Tephra fallout volume estimation from isopachs**

324 Because there was no thickness data offshore, automated contouring over the ocean cannot be  
325 constrained to obtain closed or even partial isopach coverage. Instead, five of us ('experts') drew  
326 isopach maps for the different Units, based on the same field data available on land, but extending  
327 contours offshore with each expert providing their own best estimate, judged according to their  
328 view of the field data. In most cases, they also plotted corresponding upper and lower credible  
329 bound contours to express their assessment of appropriate uncertainties associated with the  
330 contouring, given the limited data. With this methodology, the best estimate and lower and upper  
331 contours for each Unit are taken to approximate each expert's statistical median, 5<sup>th</sup> and 95<sup>th</sup>  
332 percentile contour areas, thus allowing elementary uncertainty distributions to be quantified from  
333 each expert's judgments. (Examples of Expert 1 isopachs are shown in Figure 10).

334

335 Volumes were calculated for each experts' trio of isopach maps using the Ashcalc tool (Daggitt et al  
336 2014) using the three different numerical models: Exponential (Pyle 1989), Power law (Bonadonna  
337 and Houghton 2005) and Weibull (Bonadonna and Costa 2011). We averaged the volumes  
338 determined by these three different models, with the results corresponding to the five different  
339 experts presented in Table 3. Here we note that Figure 10 shows the central "best case" contour  
340 drawn by Expert 1. Thus, the volumes indicated for Expert 1 generally fall within the range of  
341 volumes from the other experts (see Table 3) and therefore are broadly representative when  
342 compared to the other, independently drawn alternatives.

343

344 Pooling these different judgements and uncertainties to create joint volumes was treated as an  
345 expert elicitation problem. For this purpose, the Classical Model (Cooke, 1991) combinational  
346 algorithm was adopted to aggregate deposit volumes determined from the experts' maps and their  
347 uncertainties. These joint estimates were derived via the Classical Model pooling algorithm by  
348 equally weighting results obtained from individual experts' maps. On this basis, piece-wise linear  
349 distributions were computed at median and 5<sup>th</sup> and 95<sup>th</sup> percentiles for the volume of each Unit.

350

351 These combinations capture the experts' uncertainty spreads jointly, thereby synthesising the  
352 deposit volume estimates into those of a single "proxy expert", representative of the group as a  
353 whole ( n.b., the details of each deposit volume joint uncertainty distribution usually differs from  
354 those indicated by isopachs drawn by any individual expert).

355

356 Next, the piece-wise linear, equal weights Unit deposit volume uncertainty distributions -- obtained  
357 from the isopach maps as described above -- are stochastically sampled using the UNINET\* Bayes  
358 Net software package. Summing these stochastic samplings of the deposit volume uncertainty  
359 distributions, collectively, allows an overall estimate of the total volume in all the Units to be  
360 determined from the different isopachs mapped by individual experts (Mean volumes of Units 1-3  
361 range from 7.9 -24.2 x 10<sup>6</sup> m<sup>3</sup>), with related uncertainties taken into account formally (see Table 4  
362 for details). This uncertainty-based approach represents a more informative way of gauging  
363 quantitatively the eruption volumes derived from statistical methods. Reporting the eruption  
364 volumes with an uncertainty interval is preferred to using just 'best estimate' deposit volumes on  
365 their own as it can provide a clearer view of our understanding of the deposit (e.g., Constantinescu  
366 et al., 2022).

367

### 368 **Volume of Pyroclastic density currents**

369 The volume of pyroclastic density currents was estimated using the typical maximum thickness  
370 measured of the deposits in the different valleys. Thicknesses of PDCs in the Upper Wallibou valleys  
371 were not measured directly, therefore we assumed a thickness similar to the other nearby valleys of  
372 10 m. We assumed box canyon type cross sections as this is the typical present day morphology of  
373 these valleys. We obtain ~ 17 x 10<sup>6</sup> m<sup>3</sup> for the volume of the PDC deposits in the valleys outside the  
374 crater and acknowledge the considerable uncertainty in this estimate.

375

### 376 **Intracrater tephra volume**

377 Estimation of the volume of tephra inside the Summit Crater (see Table 5) was made by differencing  
378 of a post eruption Digital elevation model (DEM) made by photogrammetry in June 2021 with a pre-  
379 eruption DEM (Pleiades DEM). Almost complete destruction of the 2021 lava dome by the explosive  
380 activity means the absence of the 2021 lava dome from the pre-eruption topography does not  
381 significantly affect this calculation. This DEM differencing yields a volume of  $43 \times 10^6 \text{ m}^3$   
382 (uncertainties on this are  $\sim \pm 20\%$ , so  $\pm \sim 10 \times 10^6 \text{ m}^3$ ).

383

384 The bulk total (non-DRE) volume of tephra is  $\sim 119 \times 10^6 \text{ m}^3$  (Table 5). Lava domes form a significant  
385 contribution to the tephra deposits, we assume the 2021 lava dome had a volume of  $\sim 20 \times 10^6 \text{ m}^3$   
386 prior to the explosions and around 60% of the 1979 dome ( $\sim 28 \times 10^6 \text{ m}^3$ ) was destroyed during the  
387 explosions. Thus, the fragmented lava domes form  $\sim 40\%$  of the bulk volume of the tephra.

388 Therefore,  $71 \times 10^6 \text{ m}^3$  of the volume might be new magma-, which converts to  $42 \times 10^6 \text{ m}^3$  DRE using  
389 a 0.6 conversion based on measured tephra densities of  $1500 \text{ kgm}^{-3}$ . Here we note that the question  
390 of what constitutes “new magma” is not so straightforward as this eruption excavated and  
391 incorporated the 2020-2021 lava dome as well as incorporating new magma (see discussion below).

392

### 393 **Comparison with other volume estimates**

394 Our volume estimates are consistent with two other independent estimates using different kinds of  
395 data. Sparks et al. (this volume) used RSAM data and estimates of column heights from satellite data  
396 to estimate erupted volumes in each explosive event. Magma supply rates and magma volumes can  
397 then be inferred from these data. Their total DRE volume estimate is  $37.8 \times 10^6 \text{ m}^3$  (90% credible  
398 interval:  $[33.5 \dots 42.5] \times 10^6 \text{ m}^3$ ) which is very similar to the DRE derived from the isopach maps of the  
399 tephra deposits (DRE  $35.5 \times 10^6 \text{ m}^3$ ) using a bulk volume of  $59.1 \times 10^6 \text{ m}^3$  (Table 4). This apparently  
400 close agreement is somewhat misleading because there are uncertainties relating to proportions of  
401 non-juvenile clasts in the tephra deposits and excluding the very proximal deposits in the Summit  
402 Crater from the isopach volumes and the difficulty in estimating their contribution of heat to the  
403 plumes. These uncertainties would push volume estimates in the opposite directions, however, so to  
404 some extent cancel one another out. Camejo-Harry et al. (this volume) using ground deformation  
405 data estimate a volume change of  $50 \times 10^6 \text{ m}^3$  in the first three days of the explosive phase of the  
406 eruption and attribute this corresponds to volume (DRE) of erupted magma.

407

408

#### 409 **4. RECONSTRUCTING THE TIMING OF THE UNITS AND THEIR ASSOCIATION WITH EXPLOSIONS**

410 We use multiple strands of evidence to infer the timings of the Units within the explosive cycle, and  
411 hence the likely number of explosions or explosive pulses each unit represents. In turn, these timings  
412 can help us to reconstruct the impacts and important drivers of the changing explosive sequence.  
413 The evidence for these timings is given below and is also summarised in Table 6. We also show the  
414 timing of these Units graphically, overlain on real-time seismic amplitude measurement (RSAM) in  
415 Figure 11.

416 Re-processed GOES-16 satellite imagery that combines the IR (infrared) anomaly with ash detected  
417 in visible light ('ABI RadF ash') was used to determine the timing and distribution pattern of the  
418 explosive activity, where distinct individual explosions and matched the associated tephra. These  
419 timings correspond broadly with the seismic energy associated with each explosive pulse (expressed  
420 via RSAM, Sparks et al., this volume). Observers based on St Vincent, particularly from later on 9<sup>th</sup>  
421 April until the morning of the 11<sup>th</sup> of April, noted that individual explosive events or plumes were  
422 difficult to distinguish even during daylight hours. We have triangulated the remote sensing  
423 observations with descriptions from local observers, shared during discussion (from our first field  
424 season) or at the time of the eruption via social media, along with our field data.

425

#### 426 **U1**

427 The initial explosion of the eruption occurred at 12:41 UTC, on 9<sup>th</sup> April and was followed by a  
428 second, smaller explosion at 18:59 UTC. Photographic documentation (Fig 12a) and satellite  
429 imagery shows that from this point onwards through the night of the 9 /10<sup>th</sup> April a near continuous  
430 pulsatory plume occurred that dispersed tephra towards the ENE. The RSAM (Real Time Seismic  
431 Amplitude Measurement) shows continuous seismicity, although a number of distinct peaks can be  
432 identified (Fig 11). To be consistent with other papers in this volume and the contemporary seismic  
433 analysis (Sparks et al., this volume) we have assigned three events to this period, with a further two  
434 more distinct explosions towards the end of the U1 sequence.

435 The continuous pulsatory nature of the explosive activity from 18:59 UTC on 9<sup>th</sup> April until 04:30 on  
436 10<sup>th</sup> April is consistent with the deposit character of U1, where deposits are crudely stratified, even  
437 at relatively proximal locations. The increase in grain-size upwards (See grainsize section for more  
438 details) through U1 is associated with the latter explosions forming from higher plumes. Eyewitness

439 accounts from Windward residents on or close to the ENE dispersal axis describe pulse-like venting  
440 and continuous noise consistent with the fallout of lapilli-sized tephra on 9<sup>th</sup> overnight into 10<sup>th</sup> April.  
441 An isopach map shows that U1 had a north-easterly dispersal (Fig 10a). No fallout was reported in  
442 Barbados on 9<sup>th</sup> April, indicating that the plume dispersed to the north of the island supporting this  
443 north-easterly dispersal. The presence of ash on the ground at Calliaqua (south coast) at dawn on  
444 10<sup>th</sup> April suggests that larger U1 explosions with more widespread tephra plumes began in the early  
445 hours of 10<sup>th</sup> April.

446

447 Radar imagery (Capella Space) at 14:03 UTC on 10<sup>th</sup> April showed the presence of a crater 600m in  
448 diameter inside the southwestern portion of the old Summit Crater, with almost all of the 2021  
449 dome and ~ 60% of the 1979 dome destroyed by formation of this crater. Hereafter we refer to this  
450 as the 2021 Crater.

451

## 452 **U2**

453 The tephra plumes generated from 09:35 UTC onwards on 10<sup>th</sup> April began to form larger more  
454 axisymmetric ABI RadF ash footprint, suggesting near island-wide dispersal, and closely spaced  
455 explosive pulses, (interval  $\leq$  70 minutes) with a dispersal axis to the WSW. Photographic  
456 documentation also shows a particularly large asymmetrically spreading plume at 09:59 UTC (Fig  
457 12b). Given the timing and geometry of this explosion plume we infer that this represents the start  
458 of explosive activity responsible for U2 tephra. An isopach map for this unit shows island wide ashfall  
459 occurred being dispersed generally to the WSW (Fig 10b).

460

## 461 **U3**

462 Observers in the Leeward areas all describe an intensification of activity during the late afternoon of  
463 the 10<sup>th</sup> of April – ‘an early night’ at Spring village (Leeward) where significant ashfall was reported  
464 between evening of 10<sup>th</sup> and 11<sup>th</sup> April (with intense lightning).

465 If we assume the seven layers in U2 correspond with discrete explosive events – the seventh  
466 explosion in the U2 sequence occurred at 16:20 UTC on 10<sup>th</sup> April implying that the explosion at  
467 16:30 local time (18:30 UTC) would have been the first of the U3 tephra. However, local



468 observations are insufficiently distinct to distinguish between these two times. RSAM data suggest  
469 that the explosion at 18:30 UTC generated a more unusual RSAM peak with longer more diffuse  
470 energy and higher frequency (Fig 11) that could be PDC formation, which is associated with this  
471 phase. This later explosion is more likely to draw comments about an 'early night' and is associated  
472 with fallout in SE of the island. Thus, we suggest that the first part of the U3 sequence and the onset  
473 of PDC is associated with the 18:30 UTC explosion. The Isopach map shows a broadly axisymmetric  
474 dispersal for this unit on island with preferential dispersal to the east (Fig 10c).

475 We have ascribed three explosions to this Unit, and the third associated with the ash of U4.  
476 Atmospheric conditions at this time were wet which may explain the generation of the accretionary  
477 lapilli bearing, ash-rich deposit formed.

478

#### 479 **U5**

480 We have assigned four of the explosions to U5, although there is considerable uncertainty about  
481 this. An increase in RSAM associated with the explosion-that-cut-the-power overnight (See Fig  
482 1104:59 UTC 11<sup>th</sup> April), is consistent with an eruption that would have the intensity to generate  
483 some of the larger scoria, and we suggest that this represents some of the activity that formed the  
484 U5 unit.

#### 485 **U6**

486 By early morning on the 11<sup>th</sup> of April most observers on the Windward (eastern side of the volcano  
487 were reporting a clearing atmosphere and very little further ash fell. Thus, we are confident that this  
488 is the time when it is most likely that the accretionary lapilli bearing ash layers of U6, which are  
489 abundant on the Western side of the volcano was formed from a sequence of explosions. U6  
490 accretionary lapilli bearing ash layers are interbedded with PDC deposits in the Larikai and Roseau  
491 valleys, indicating that PDCs continued be formed in those valleys, associated with these explosions  
492 from 10: 36 UTC on 11<sup>th</sup> April.

493 The presence of accretionary lapilli is testament to available moisture in the plumes that formed  
494 these units. As they the accretionary lapilli –rich ash layers of this Unit are mainly seen in deposits  
495 on the south-west side of the volcano (Fig 10b) one possibility is that the ash layers associated with  
496 this Unit were formed as a result of PDCs entering the sea in the Larikai and Roseau valleys.

497

498

499 **U7**

500 Following the 11 April activity, explosions had a lower inferred column height (and a lengthening  
501 repose between explosion interval) consistent with the deposits of U7 chiefly being observed on the  
502 Leeward side of the island (Fig 3) with the greater part of the ash column being transported by the  
503 Tropospheric trade winds.

504

505 **Timing and duration of PDCs**

506 Field evidence indicates that the first PDCs to travel out of the Summit Crater were associated with  
507 U3 explosions. Satellite imagery and eyewitness observations indicates that U3 commenced before  
508 the end of 10<sup>th</sup> April, perhaps in the later afternoon/ early evening (see table 6 and Fig 11). Partially  
509 carbonised trees washed up at Chateaubelair at midday on 11<sup>th</sup> April also support the occurrence of  
510 PDCs earlier that day or on the evening of the 10<sup>th</sup> April. Evidence also indicates that U5 activity  
511 began in the early hours of 11<sup>th</sup> April and PDCs in the proximal SE flank were clearly associated with  
512 these explosions. Radar imagery (Disaster charter imagery) acquired on 11<sup>th</sup> April 19:56 showed that  
513 PDCs had entered several valleys to the south and south-west before that time.

514 Two lines of field evidence clearly demonstrate that PDCs were associated with several discrete  
515 explosions, possibly over a number of days: ash fallout layers are interbedded between some PDC  
516 flow units in both the Larikai and Roseau valleys; and secondly fallout layers U3, U4 and U5 are  
517 interbedded between PDC flow units in the proximal south-eastern area. Although evidence  
518 indicates that PDCs continued to be formed associated with U6 and U7 and were probably  
519 concentrated in the Larikai and Roseau valleys.

520

521 **5. DISCUSSION**

522 The initial explosive activity between 12:36 UTC on 9<sup>th</sup> April and 9:35 UTC on 10<sup>th</sup> April resulted in  
523 the destruction of the 2021 lava coulee and ~60% of the 1979 lava dome, and formation of the 600  
524 m wide 2021 crater that was observed in radar imagery at 14:03 UTC on 10<sup>th</sup> April. This activity  
525 formed the U1 lapilli deposit that is ash poor and rich in hydrothermally-altered, semi-vesicular and  
526 non-vesicular material. It is likely that some of these dense fragments have been sourced from  
527 these pre-existing lavas. Crude layering and reverse grading present in U1 together with an increase

528 in grain size and abundance of vesicular scoria, testifies to the pulsatory nature and increasing  
529 intensity and our grain type analysis to an increasing proportion of primary vesiculating magma.  
530 These observations are consistent with the early stages of explosive eruptions in which near surface  
531 overpressures are high, resulting in excavation of a 2021 crater and conduit system (Sparks et al.  
532 2006).

533 After the initial explosive phase that formed U1 deposits and the 2021 crater, there was a pause in  
534 explosive activity for more than four hours during which a swarm of VT earthquakes occurred, with  
535 some evidence that these became deeper with time (Contreras-Arratia pers comm). These coincided  
536 with rapid deflation inferred from surface deformation data (Camejo-Harry et al., this volume) and a  
537 pulse of SO<sub>2</sub> associated with the initial plume (Esse et al., this volume). We propose that these VTs  
538 were associated with the rapid ascent of a new pulse of gas-rich magma that resulted in the  
539 character of the deposits associated with Units 2 – 5.

540 Furthermore, our analysis of bubble density and shape (Fig 9) suggests that vesiculation processes  
541 were inhibited by the microlite-rich character of the magma at that time, as well as the work to be  
542 done excavating the 2021 crater and removing the overlying lava. We interpret the sequence of  
543 tephra deposits in terms of a model developed for kimberlite explosive volcanism (Sparks et al.  
544 2006) but which is more generally applicable. In this model explosions in an initial overpressured  
545 regime excavate a crater and conduit followed by a pressure-adjusted regime which includes infilling  
546 of the crater-conduit systems with tephra and clastic materials derived from further instability of the  
547 2021 crater and conduit walls.

548 Juvenile magma from U1 shows a notably lower vesicularity than that in U2 (Fig 8 and Table 2).  
549 Microtextural studies (Frey et al, this volume) show that U1 juvenile material was distinctly richer in  
550 microlites, and contained different phases, than later Units. The U1 magma was likely from a  
551 different source to those of later Units. Texturally it is the most distinctive of our erupted magmas,  
552 the groundmass is dominated by crystals with a short axis < 5 µm and aspect ratios greater than 1 :  
553 3, median vesicle area is the smallest (109 µm<sup>2</sup>) and it has the lowest BND (167 bubbles/mm<sup>2</sup>).

554 Activity from 9:30 to 18:50 UTC on 10<sup>th</sup> April was characterised by closely spaced explosions (<1 – 2  
555 hrs apart) and is correlated with the ash-rich U2 deposits. We attribute the change to the influence  
556 of an established crater-conduit system in which material falling back into the conduit interacts with  
557 erupting magma. The 2021 crater with a tephra rim of up to 40 m thick had formed early part of U2  
558 activity. We envisage instability of the unconsolidated tephra rim and crater-conduit walls together

559 with fall back of tephra into the 2021 crater and conduit led to generation of ash through  
560 fragmentation of these recycled materials.

561 Field evidence strongly indicates that onset of the first significant PDC activity was associated with  
562 U3 activity from 21:30 UTC on 10<sup>th</sup> April. Proximal tephra sequences show no evidence of PDCs until  
563 after deposition of ash-rich U2 tephra fallout. Distinctive U3 lapilli fallout interbedded with PDCs  
564 testifies to the onset of PDCs at this time. In addition, particularly large axisymmetric explosion  
565 derived plumes causing early darkness in the afternoon of 10<sup>th</sup> April is coincident with vigorous  
566 explosions of U3. We suggest that processes, such as conduit widening (or perhaps when the  
567 excavation of looser tephra was more complete) are likely causes of this transition to PDC activity,  
568 this seems likely as > 24 hours of semi continuous explosive activity had already occurred. As U3  
569 fallout deposits are some of the coarsest, most vesicular of the sequence, we can rule out a  
570 decreasing gas content of the magma playing an important role in the switch to collapsing columns.

571 The U3 activity also marks the onset of the appearance of aggregates and accretionary lapilli within  
572 the tephra fall deposits. Accretionary lapilli and ash aggregation only need moisture-rich plumes to  
573 form. There are several sources of moisture which could have promoted ash aggregation. First,  
574 moisture-rich weather systems generating rainfall on the island associated with the explosive activity  
575 occurred from 11<sup>th</sup> April. Another origin for the moisture is from PDCs entering the sea. During the  
576 1995-2010 activity at Soufriere Hills, Montserrat accretionary lapilli were formed at numerous times  
577 from PDCs entering the sea and generating fallout layers rich in these ash aggregates (Burns et al.  
578 2017). Finally, it is possible that ground waters ingressed into the conduit but there is no evidence of  
579 hydrovolcanism so we think this is a less likely mechanism. We therefore consider that these  
580 accretionary lapilli were formed where either tephra plumes interacted with moisture-rich weather  
581 systems or where moisture-rich plumes were formed by PDCs that entered the sea.

582 Accretionary lapilli are abundant in numerous ash layers within the upper part of the sequence.  
583 Some of these ash layers are likely formed from elutriation of ash from PDCs, and the abundance of  
584 ash layers in sequences on the southwestern flank supports a co-PDC origin. The ash layer that  
585 occurs between the two coarse lapilli layers of U3 and the U4 ash layer are both likely to have been  
586 co-PDC ash fallout layers, occasional coarse clasts embedded within the U4 ash probably relate to  
587 continued minor explosive activity during settling of these ash layers.

588 The implications from the distribution and size of the scoria associated with the U5 explosions are  
589 that these were some of the largest and strongest plumes during the course of eruption. This  
590 requires some explanation as at this time the explosions had been ongoing for around 40 hours, and

591 are close to the point at which Sparks et al., (this volume) have identified a waning stage for the  
592 eruption based on seismic energy and plume height. These scoria have a normal distribution of bulk  
593 vesicularities, and the lowest bubble number density and greatest vesicle areas (77 bubbles/mm<sup>2</sup>).  
594 With both U2 and U3 scoria showing wider distributions of vesicularities, we infer from this that U2  
595 and U3 explosions were composed of a mixture of the initial magmas as well as driven from behind  
596 by the newer magmatic batch, a concept supported by the changing microlite populations identified  
597 in Frey *et al.*, (this volume). We postulate that at the point of the U5 explosions the first source was  
598 exhausted and U5 explosions are only driven by the comparatively hotter (Frey et al., this volume)  
599 new magmatic batch. The U6 and U7 explosions account for the slow depletion of this magma batch  
600 and increasing repose periods between explosions until insufficient momentum remained to  
601 generate explosions. It also coincides with the period of rapidly declining magma supply rate (Sparks  
602 et al., this volume)

603

#### 604 **The character and temperatures of PDCs**

605 It is evident that PDCs were generated by collapsing eruptions columns or fountains. Similar to those  
606 formed in other eruptions of this volcano (Hay 1959; Cole et al. 2019). Those that reached the sea at  
607 the mouths of the Larikai and Roseau valleys were by that point likely moving at low velocities, as  
608 there is no associated dilute PDC component in that region. Dilute PDCs only extended to a  
609 maximum distance of 2 km, inundating ridges to the south and west.

610 Field evidence also indicates that dilute PDCs in the proximal regions (< 2 km) had temperatures too  
611 low to carbonise vegetation (<250° C). Those PDCs which moved down the Larikai and Roseau valleys  
612 were hotter, with temperatures high enough to carbonise most wood fragments incorporated  
613 (probably > 300° C).

614 One possibility for these temperature differences is that the PDCs which reached the sea were  
615 formed from different explosions that had differing temperatures. A second possibility is that the  
616 dilute PDCs entrained significant quantities of air, enough to cool that portion of the PDCs below the  
617 carbonisation temperature of ~ 200° C , whereas the dense PDCs that moved down the valleys  
618 retained enough heat to carbonise the vegetation incorporated within them. A third explanation is  
619 that the duration of the dilute PDCs was brief, perhaps of the order of a few minutes, not long  
620 enough to result in carbonisation of the trees, whereas those incorporated in valley-confined PDCs  
621 remained engulfed in the hot deposits and were therefore carbonised.

622

623 **Comparison to other eruptions at La Soufriere St Vincent and elsewhere**

624 La Soufriere St Vincent is the most active terrestrial volcano in the Eastern Caribbean. Prior to 2021,  
625 there had been at least six explosive eruptions over the last 1000 years (Cole et al 2019). It is  
626 therefore of considerable interest how the 2020/2021 eruption compares to these previous  
627 eruptions.

628 Our volume estimations indicate that the 2021 explosive eruption was of VEI 4 magnitude. The 1979  
629 was considered to be VEI 3, although the PDCs deposits have a relatively similar distribution at least  
630 in the Roseau and Larikai valleys (Sigurdsson et al. 1979). Tephra fallout on island associated with  
631 the 1979 eruption is notably thinner (Brazier et al. 1982). Although no detailed isopach maps have  
632 been constructed for the 1902 eruption) the extensive, radial nature of PDCs around the volcano,  
633 and the work of Pyle et al (2018) indicate that the combined output during the 1902 eruptions on  
634 the 6-7<sup>th</sup> of May would put this eruption VEI 5 category. The 1902-03 eruptions were perhaps  
635 unique in that following the intense activity over 24 hours on the 6<sup>th</sup>-7<sup>th</sup> of May 1902 there were  
636 subsequently a further 4 further discrete explosive episodes, the last of which (in March 1903)  
637 produced basaltic scoria (Cole et al., 2019).

638 The 1812 eruption lasted for several days from the 27<sup>th</sup> of April to the 1<sup>st</sup> of May with a residual  
639 explosion on the 6<sup>th</sup> May (Hugh Perry Keane Accounts, quoted in Smith, 2011). From contemporary  
640 accounts the most violent phase was from the 30<sup>th</sup> April to the 1<sup>st</sup> of May and although nearby  
641 inhabitants were engulfed in darkness and pelted with 'showers of stones & earth and rocks' (Hugh  
642 Perry Keane in Smith, 2011) it is clear from these accounts that this consisted of a repetitive series of  
643 explosions. Accounts indicate that PDCs generated occurred only after several days of explosive  
644 activity that generated tephra fallout (Shepherd 1831 and contemporary Parliamentary Record). This  
645 pattern is similar to the 2021 explosive activity, where PDCs only occurred following > 24 hours of  
646 significant explosive activity. However, our preliminary geological observations indicate that these  
647 1812 PDCs, although not as extensive as 1902, were considerably larger than those of the 2021  
648 events, and, as yet there is no written evidence for a preceding dome-forming eruption.

649 Prior to both the 1902 and 1979 a lake was present in the Summit Crater and indeed for both these  
650 events hydrovolcanic activity has been invoked for the eruption mechanism of the 1979 explosions  
651 (Shepherd et al. 1979; Shepherd and Sigurdsson, 1982), and owing to abundant lithic material, for  
652 the opening phases on the 1902 eruption (Cole et al. 2019). The 2020/2021 eruption differs in that it  
653 began with three months of lava dome emplacement, and there was no evidence of hydrothermal  
654 water involvement or steam-rich plumes during the initial explosive phases. Thus, it seems the 2021

655 explosive eruption falls between 1979 and 1902 in terms of magnitude, but was quite similar in  
656 pattern and style to the eruption in 1812.

657 An important feature of the main phase of the 1902 eruption, the accounts of the 1812 eruptions  
658 and the 2021 eruption has been the presence of incessant and repetitive explosions that at some  
659 point during the eruption also generated PDCs and lahars. Our initial analysis of these deposits prior  
660 to their substantial erosion provides a unique record of subtle changes that lead to important  
661 changes in eruptive behaviour and impacts, paving the way for more detail study. Comparison of  
662 these within the context of the other deposits could provide insights into the critical drivers of  
663 eruptive transitions and these repetitive sequences of volcanic explosions.

664 Similar recent eruptions include Kelud in Java, Indonesia, where a short-lived explosive eruption  
665 occurred in 2014 of VEI 4 magnitude, generating both PDCs and extensive tephra fallout (Maeno et  
666 al. 2019). Tephra fallout generated by the 2008 eruption of Chaiten, Chile despite being rhyolitic, is  
667 similar, with multiple layered tephra fallout deposit formed by different phases of the eruption  
668 (Alfano et al 2011 and 2012) and associated PDCs (Major et al. 2013). Although Chaiten and Kelud  
669 generated blast-like PDCs, for which there is little evidence at La Soufriere St Vincent.

670 In summary, the 2021 eruptions of La Soufrière presented serious hazards to the population  
671 (~20,000) living around the volcano. Although PDCs generated were limited, in that they reached the  
672 coast only in two valleys to the west, similar to 1979, only a slight increase in magnitude and  
673 intensity of future eruptions would result in more extensive and longer runout PDCs as happened in  
674 1902.

## 675 **Conclusions**

- 676 1. Explosive activity between 9<sup>th</sup> and 22<sup>nd</sup> April 2021 resulted in a layered tephra deposit. The  
677 lowermost deposit, Unit 1 was associated with destruction of almost all of the 2021 lava  
678 dome and ~60% of the 1979 lava dome resulting in a 600m diameter 2021 crater.  
679
- 680 2. Vesicular scoria representing the juvenile magma is present throughout the tephra  
681 sequence, but in varying quantities. Documented variations in the nature of the vesicularity  
682 is consistent with different batches of magma rising through the system in the first few days.  
683
- 684 3. PDCs were only formed after > 24 hours of explosive activity, contemporaneous with several  
685 explosions. The switch from convecting to collapsing plumes is inferred to be related to vent  
686 processes, possibly flaring/widening. Dense, concentrated PDCs reached the sea in the

687 Larikai and Roseau valleys only. Dilute PDCs were limited to within 2.5 km of the Summit  
688 Crater rim but were distributed across the lower crater rim from west to southeast.

689

690 4. Volume calculations indicate that the explosive phase of this eruption had a bulk volume of  
691  $1.19 \times 10^8 \text{ m}^3$  ( $7.1 \times 10^7 \text{ m}^3$  DRE) +/- approximately 20% and therefore was a small VEI 4  
692 eruption. Approximately one third of the volume of tephra is contained within the pre-  
693 existing Summit Crater forming deposits locally > 100m thick.

694

695 5. The changing style and characteristics of the tephra deposits can be attributed to changes in  
696 the ascending magma and the evolution of the conduit-crater system during the explosive  
697 eruptions. In the first 20 hours a new explosion crater and deeper conduit formed.  
698 Thereafter back-filling of the 2021 crater-conduit system by crater wall rock instability and  
699 recycling of tephra led to interactions of erupting magma with back-fill resulting in more ash-  
700 rich deposits and conditions for column collapse and PDC formation.

701

702 6. Although ash aggregates such as accretionary lapilli were abundant in the upper half of the  
703 tephra sequence, there is no evidence for hydrovolcanism, and we infer moisture was  
704 derived from weather systems or PDCs entering the sea.

705

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713

714

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816

## 817 **Figures**

818 Fig 1 The island of St Vincent and location of La Soufriere volcano on the island. Inset are the islands  
819 of the Eastern Caribbean arc. The area shaded red represents the volcanic edifice.

820 Fig 2 Measured sections through the 2021 tephra sequence Leeward sections (top) and Windward  
821 locations (bottom).

822

823 Fig 3 – Tephra sequence at a) Owia (5 km NE) scale in cm divisions, b) Dry Wallibou mouth (4.5 km  
824 SW of Summit Crater rim), c) River Bed 1.8 km SE, Sequence is 30 cm thick in total. d) Tephra fallout  
825 sequence at ‘The Bench’ 1.2 km SE of Summit Crater rim, sequence is a total of 45 cm thick. e)  
826 Sequence at Loc 38, Dark View falls 6 km SW of Summit Crater rim. Sequence is 3.8 cm thick. Scale is  
827 in 1 cm intervals. f) Accretionary lapilli in Unit 4 at Jacobs well (700 m SE of Summit Crater rim). Coin  
828 is 2cm across.

829 Fig 4 a) Map showing extent of PDCs, from fieldwork and satellite imagery. Red =dense valley filling  
830 PDCs. Pink = dilute PDC; PDC deposits. Yellow dots are sections shown in b-f; b) in the Larikai, c)  
831 Roseau, d) section at Jacobs well showing PDCs interbedded with fallout layers, e) Upper SE flank  
832 valley (looking South from Jacobs Well) f) Several PDC flow units interbedded between with fallout  
833 layers at Jacobs well (Upper SE flank shown in a) and d)

834

835 Fig 5 a) Tree felling ~1km SW of the Summit Crater rim on the Leeward trail. PDCs moved from left to  
836 right. Note several trees are broken 2-3 m above their base. b) Detail of trees impacted by PDCs in  
837 same region as ‘a’. PDCs moved from right to left obliquely into the plane of the photograph. Note-  
838 abrasion of trees on volcano facing (RHS) side and absence of charring of trees. c) Deposits of dilute  
839 PDCs in a similar region to a) showing strong lateral thickness variations and crude cross-bedding

840 Current direction is from left to right. Spade handle rests on U2. Pen rests on lowermost PDC  
841 deposits and is 14 cm long.

842

843 Fig 6 a) Cumulative grainsize distributions for the different units. U1 and U2 are shown as areas.  
844 Units 3 to 5 are depicted as lines for individual grainsize analyses b) cumulative grainsize  
845 distributions of the different lithofacies Lapilli and Ash layers shown as shaded areas. PDCs (dilute  
846 and concentrated) are plotted as lines of individual grainsizes analyses. Grainsize histograms and  
847 statistics through two sequences C) Windward: Mahan ridge and Sandy bay d) Proximal – Jacobs well  
848 (700 m SE). Closed symbols on Max clast sizes are dense clasts, open symbols= vesicular scoria.

849 Fig 7 a-d SEM images of component types a) dense b) semi vesicular c) vesicular scoria d)  
850 hydrothermally altered. e) Variations in component types > 1mm through three sequences of 2021  
851 tephra on the Eastern (Windward) side of the volcano.

852

853 Fig 8 Vesicularity variations in juvenile clasts 16 - 32 mm in four different Units. Clast bulk densities  
854 were calculated using the Archimedes principle after Shea et al., 2010 and converted to vesicularity  
855 using a DRE value of 2.79 g/cm<sup>3</sup> – as calculated using whole rock (XRF) data and the method of  
856 Lange and Carmichael (1987) In brackets, n = number of clasts analysed from each unit. Pie charts  
857 show the proportion of field classified vesicular and semi-vesicular clasts in the different measured  
858 vesicularity ranges.

859

860 Figure 9: Density plots of bubble shape and size from Units 1, 2 and 5 (A – F), dashed red line marks  
861 boundary at which roundness = circularity. Backscattered electron images showing changing bubble  
862 textures between Units 1, 2 and 5 (G – I) and Unit 3 (J and K). SV = clast categorised as semi-  
863 vesicular, S = clast categorised as scoriaceous. Numbers in brackets refer to 3D bulk vesicularities  
864 acquired using the Archimedes method following Shea *et al.*, (2010). *Circularity* is sensitive to  
865 deviations from a perfect circle via either surface roughness (e.g., protrusions caused by crystal  
866 interference or preservation of coalescence features), or elongation. In contrast, *Roundness* is a  
867 measure is less sensitive to protrusions on vesicle walls (Liu *et al.*, 2015). If elongation is the  
868 dominant factor, lower values of *Circularity* will be approximately equal to *Roundness* (Liu *et al.*,  
869 2015)

870

871 Fig 10 Hand-drawn isopach maps for Expert 1 for a) U1, b) U2, c) U3, d) U4-7, e) total tephra  
872 thickness and f) interpreted distal thickness including Barbados island. Thicknesses values are in  
873 millimetres.

874

875 Fig 11 Timeline showing Real time seismic amplitude (RSAM) for (a) the whole explosive phase and  
876 (b) a detailed of RSAM and seismic frequency spectra for the first 100 hours of explosive activity. Our  
877 interpretation of the timing of the different Units (see also table 6) is superimposed on this, start  
878 times of each Unit given in Local time and UTC in parentheses. U4 is not shown simply owing to  
879 space issues. Fig 12 a) Tephra plume at 5:15 LT (21:15 UTC) on 9<sup>th</sup> April 2021 viewed from boat just  
880 north of Kingstown approximately 19 km southwest of the volcano. Photo credit: Kai Best b) Large  
881 axisymmetric spreading plume at 5:59 LT (09:59 UTC) viewed from nr Calliaqua, close to the  
882 southernmost point of the island, 22.5 km south of the volcano. This plume likely corresponds to the  
883 start of the Unit 2 phase. Photo Credit: Jenny Trumble

#### 884 **Tables**

885 Table 1 Summary of key features of different Units identified. N.G=normal grading. R.G. = reverse  
886 grading. S= vesicular scoria, SV= semi vesicular, D= Dense, H = hydrothermal

887 Table 2: Key measurements from 2D image analysis of vesicles and micro/phenocryst phases for  
888 Units 1, 2 and 5.

889

890 Table 3 Bulk volumes of the different Units calculated from isopach maps. Volumes in millions of  
891 cubic metres (to the nearest million). 'Best Case' = in bold, 95 and 5% estimates in parentheses

892

893 Table 4 Results of UNINET stochastic sampling of the EqWt Unit bulk volume piece-wise linear  
894 uncertainty distributions from EXCALIBUR. Volumes are  $10^6\text{m}^3$

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896 Table 5 Bulk volumes of the different components of the explosive activity.

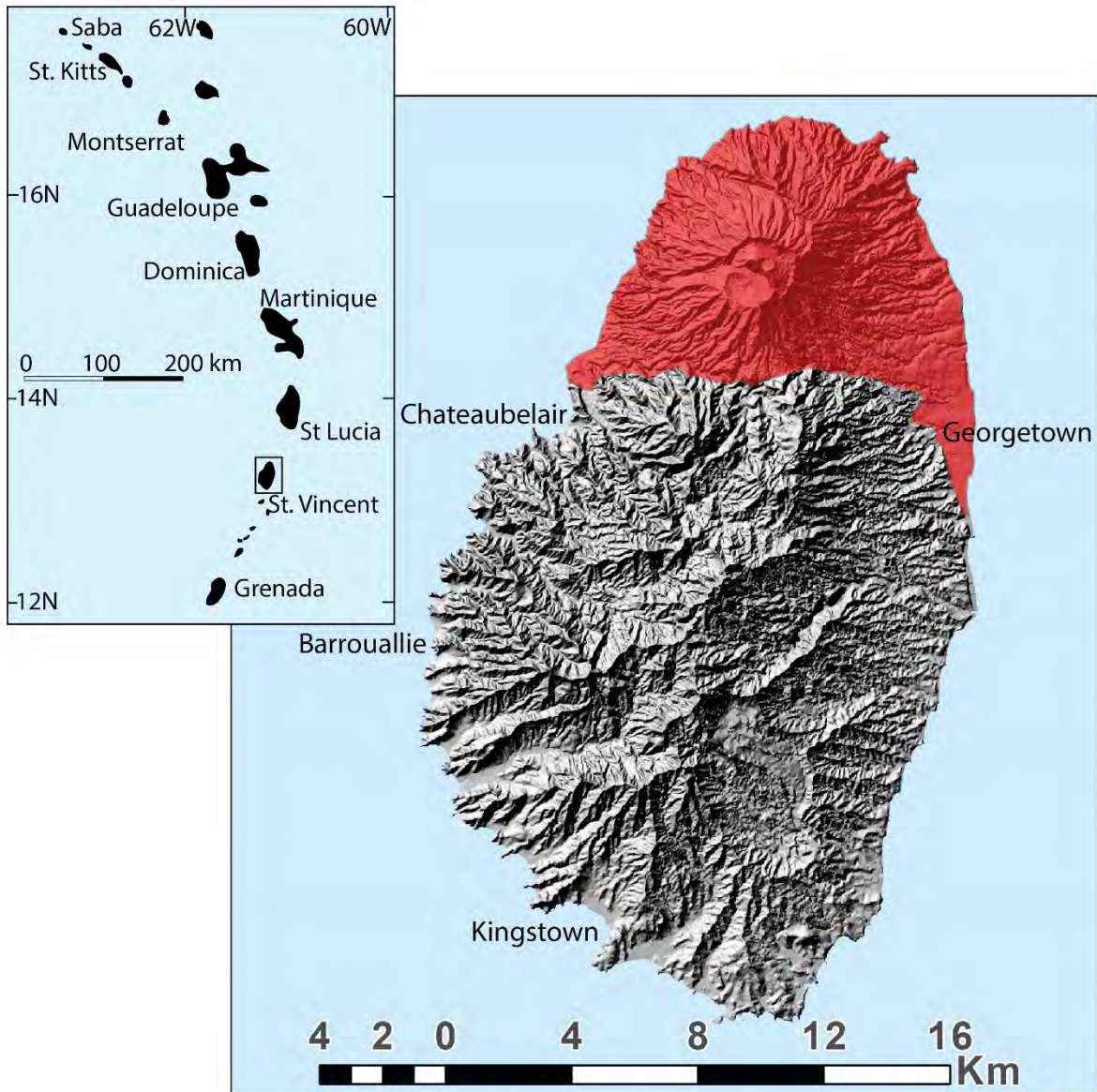
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898 Table 6: (1) Explosion number – explosion assignment from Robertson et al., this volume (2) Seismic  
899 (RSAM data) to define onset on explosions and durations of RSAM signal. Data used here are RSAM  
900 Spike determinations by Sparks et al., this volume but see also Fig 10. (3) Times used here are  
901 derived from satellite data as described in Sparks et al., (this volume). Dark grey shading indicated  
902 discrete explosions less than two hours apart, light grey from 2-4 hours apart. Thick black line  
903 indicates high degree of certainty on boundary timing, lighter grey lines indicate lower degrees of

904 certainty.

905

906 **Fig 1**



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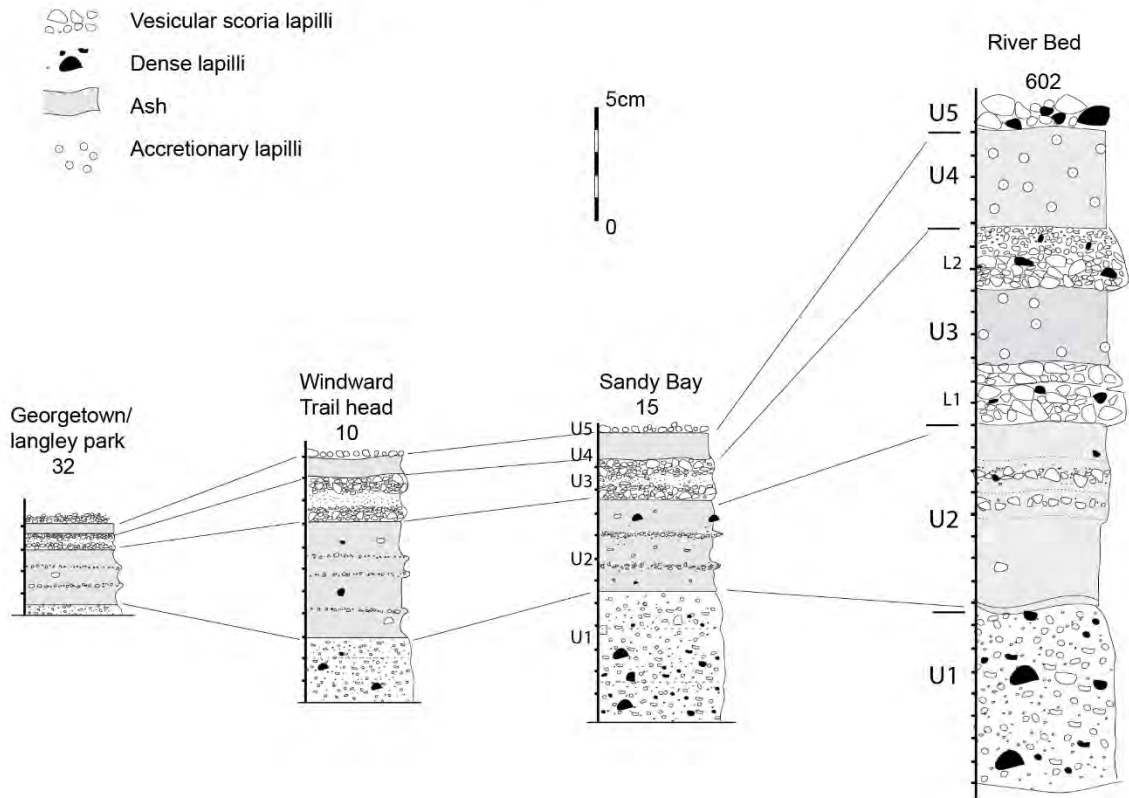
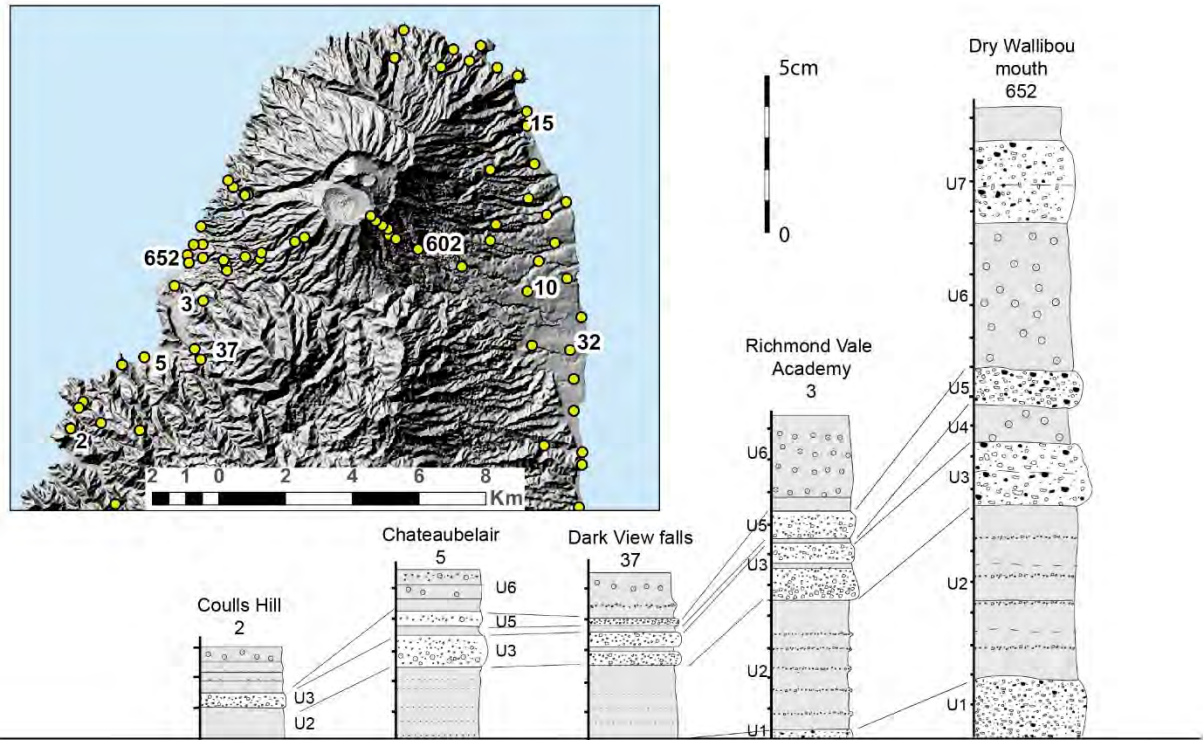
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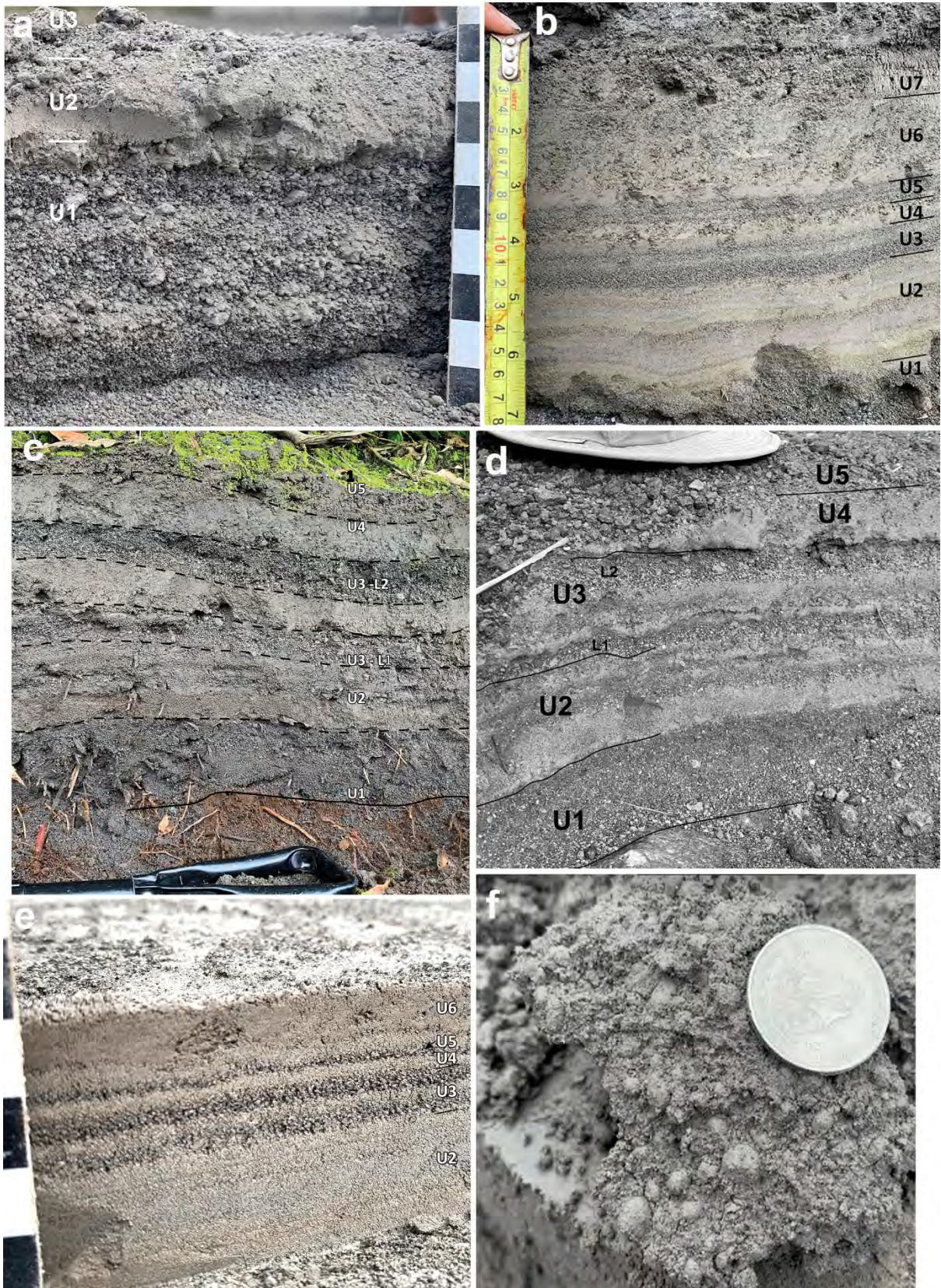
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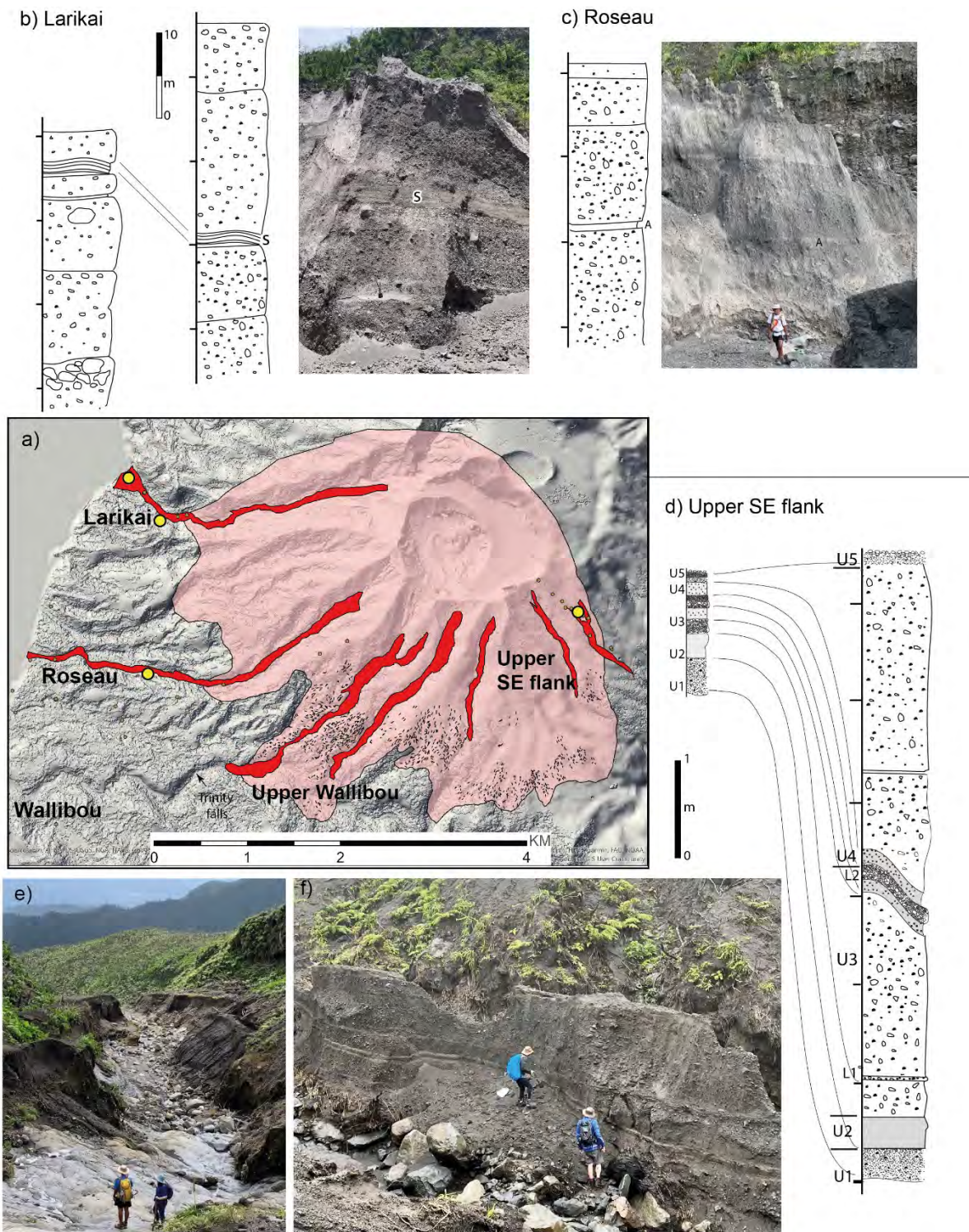
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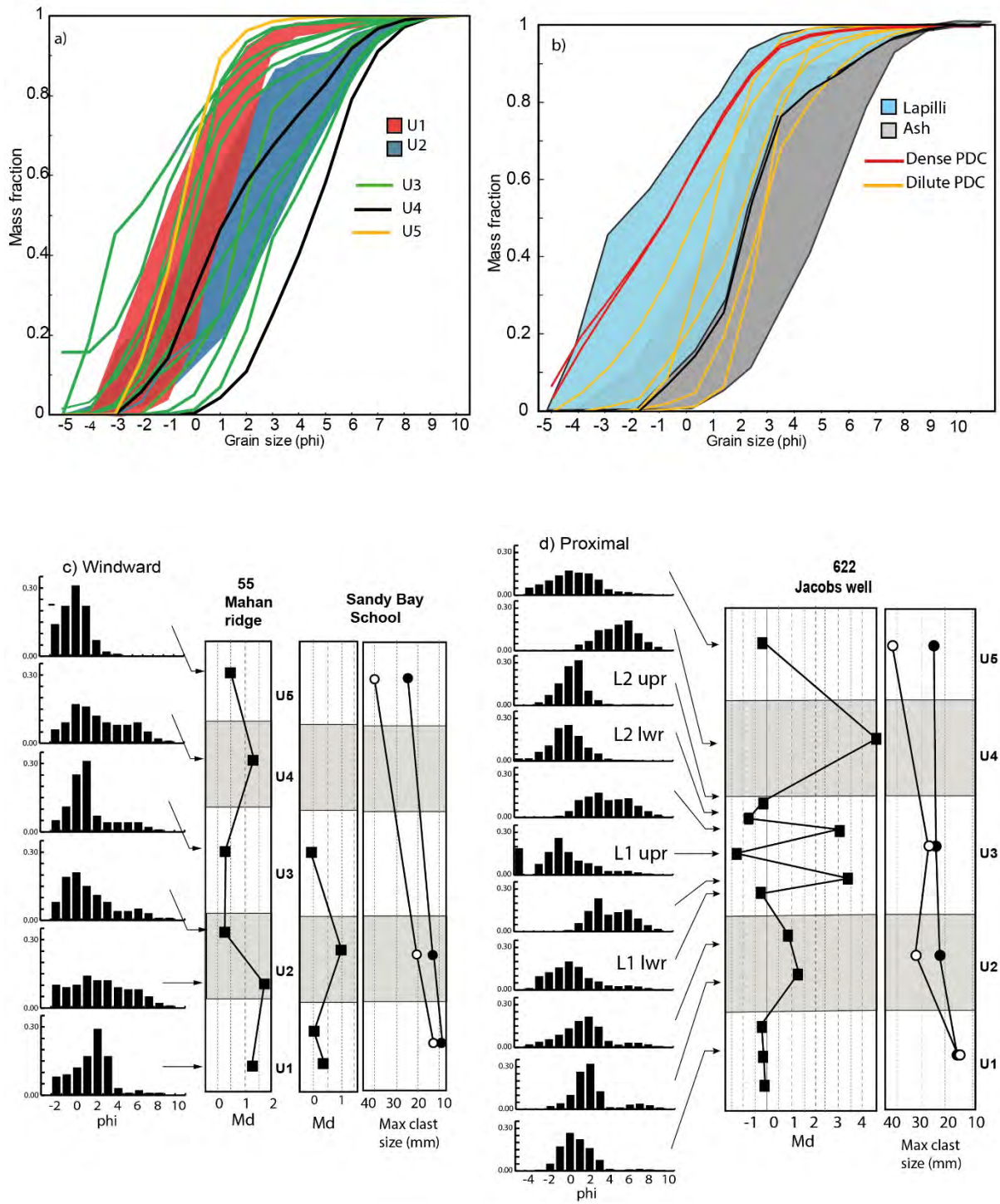
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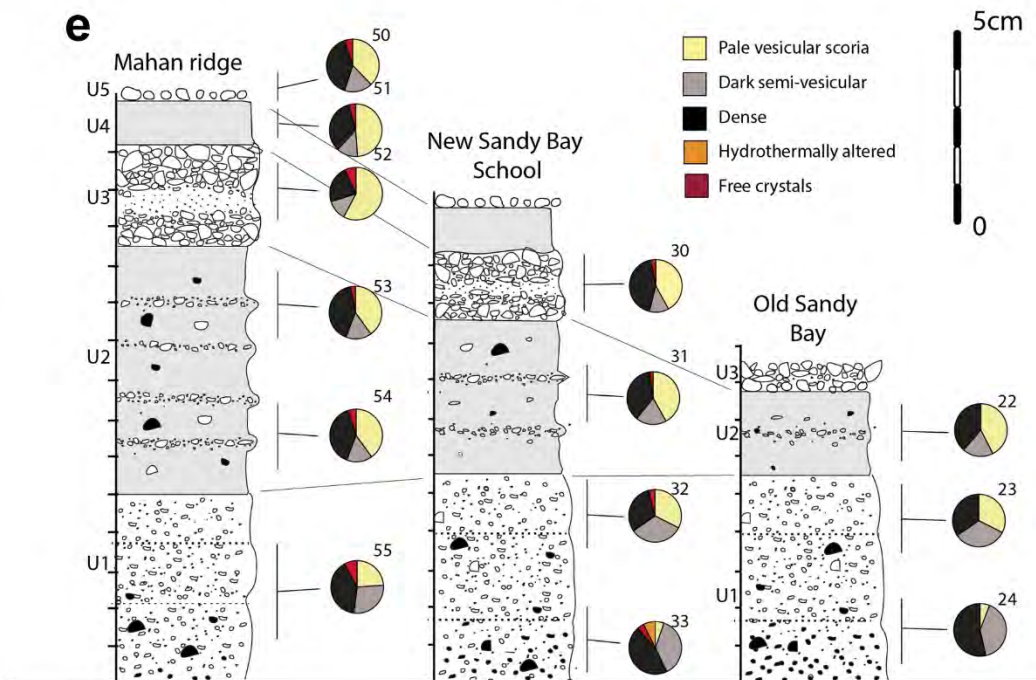
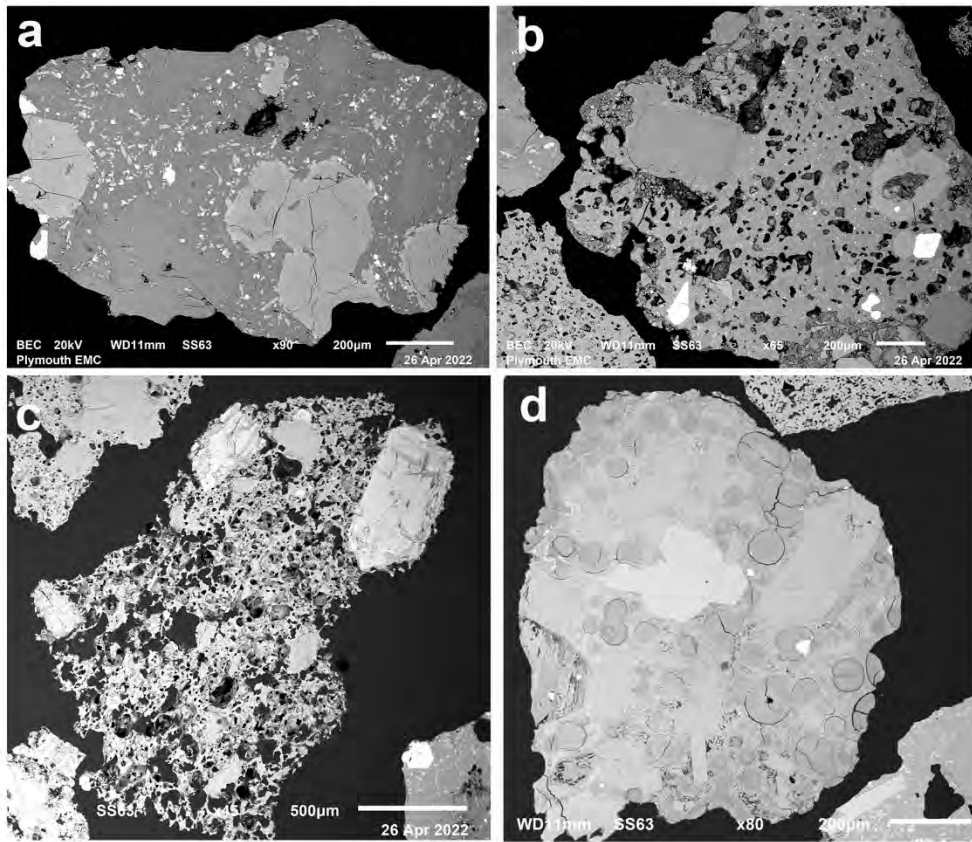
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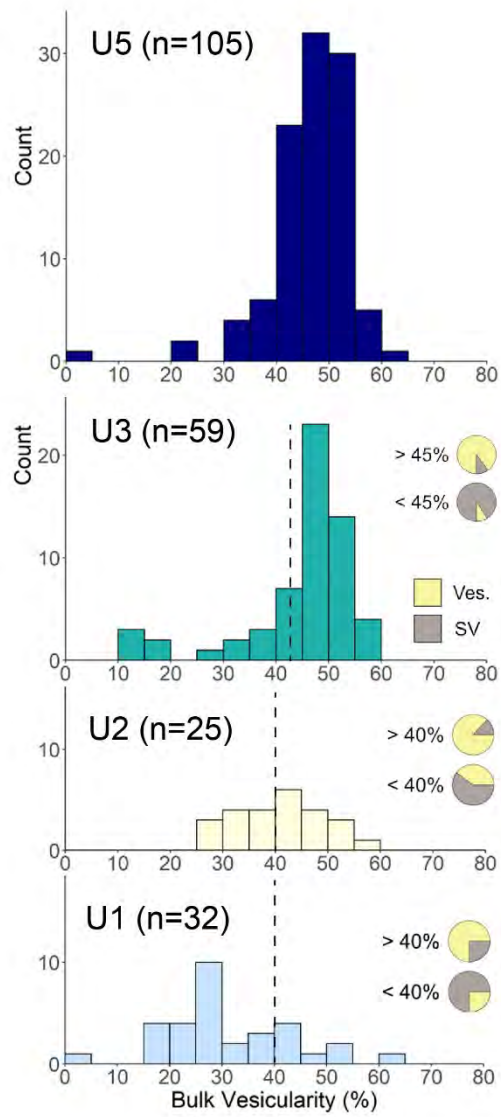
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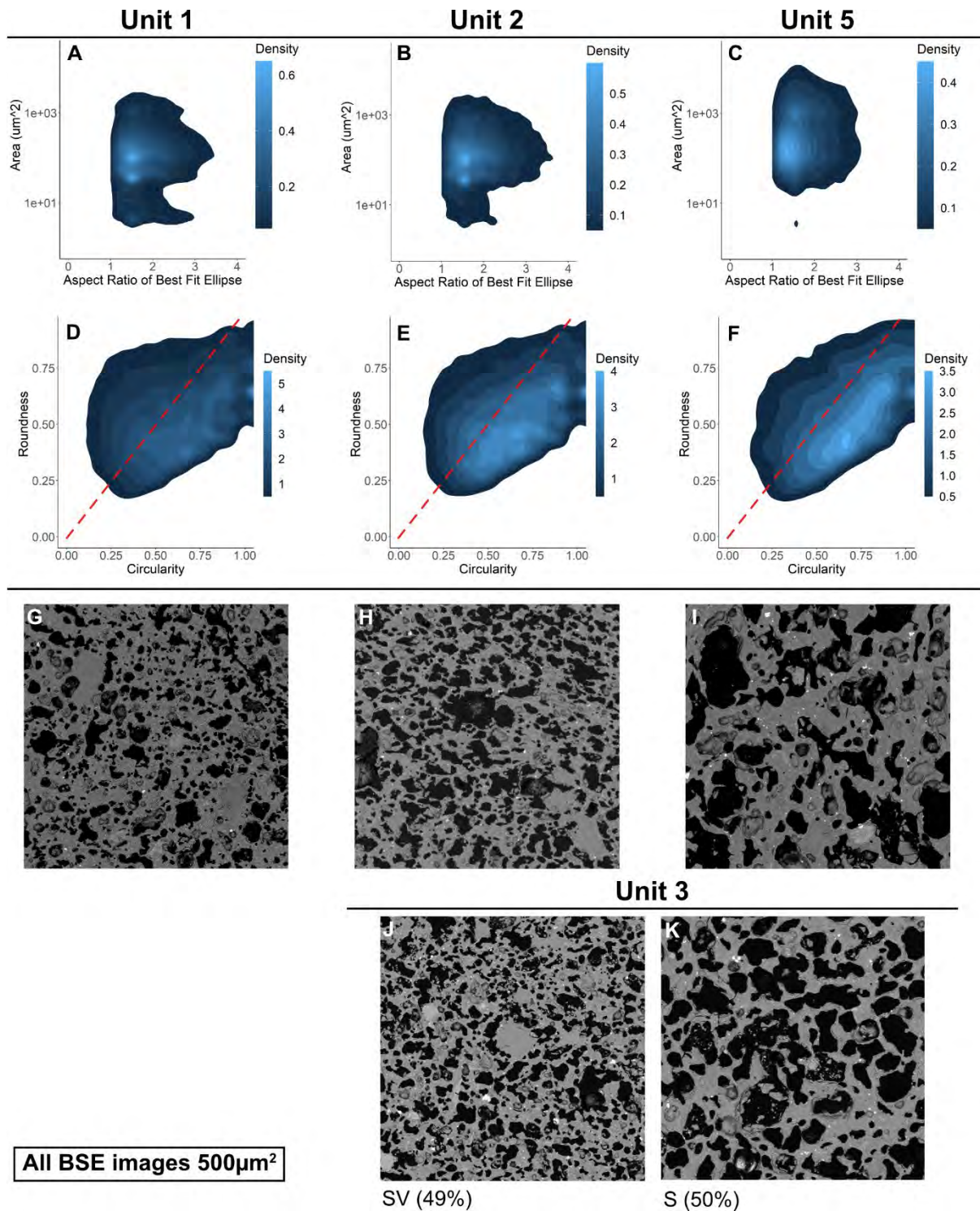
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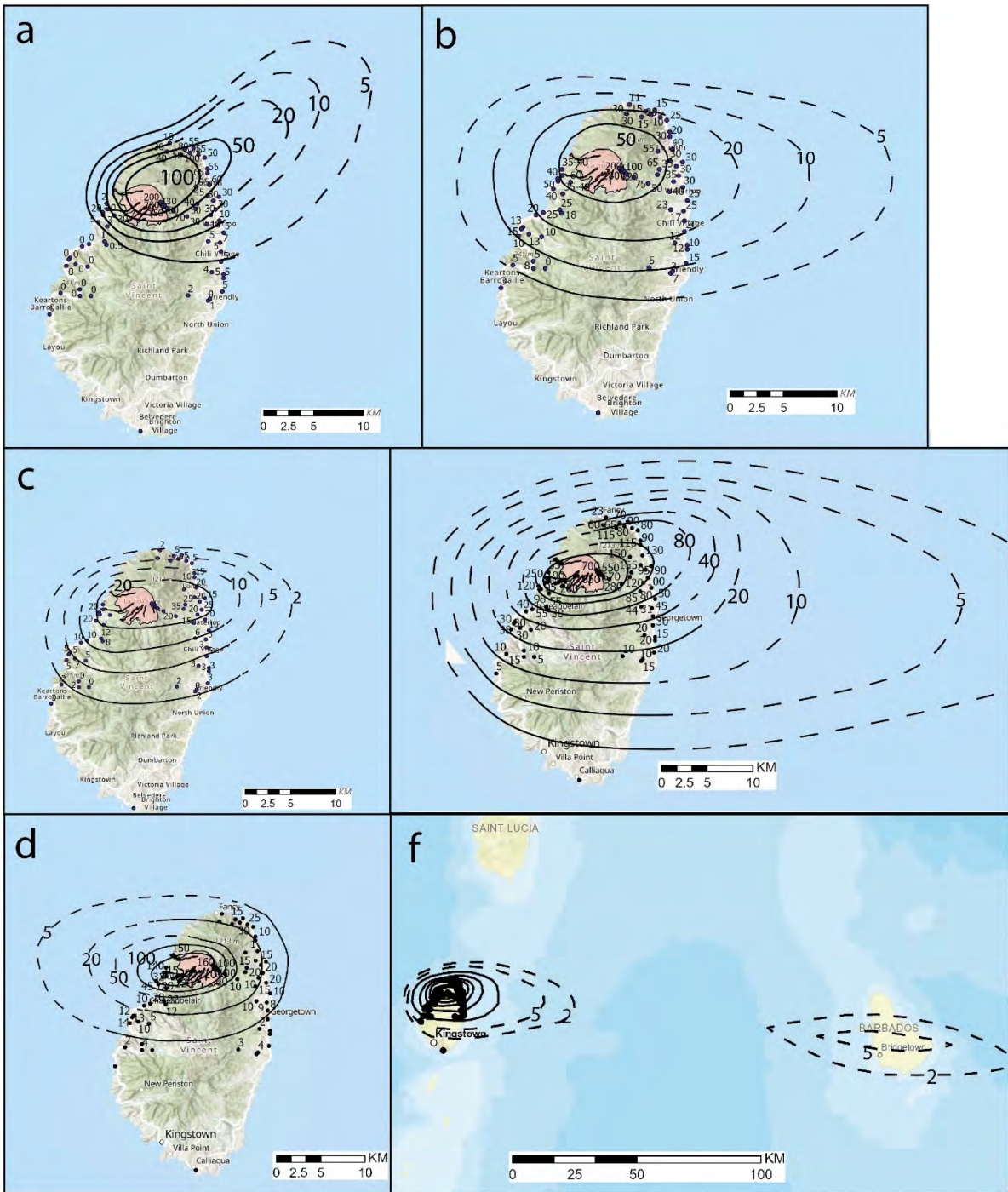
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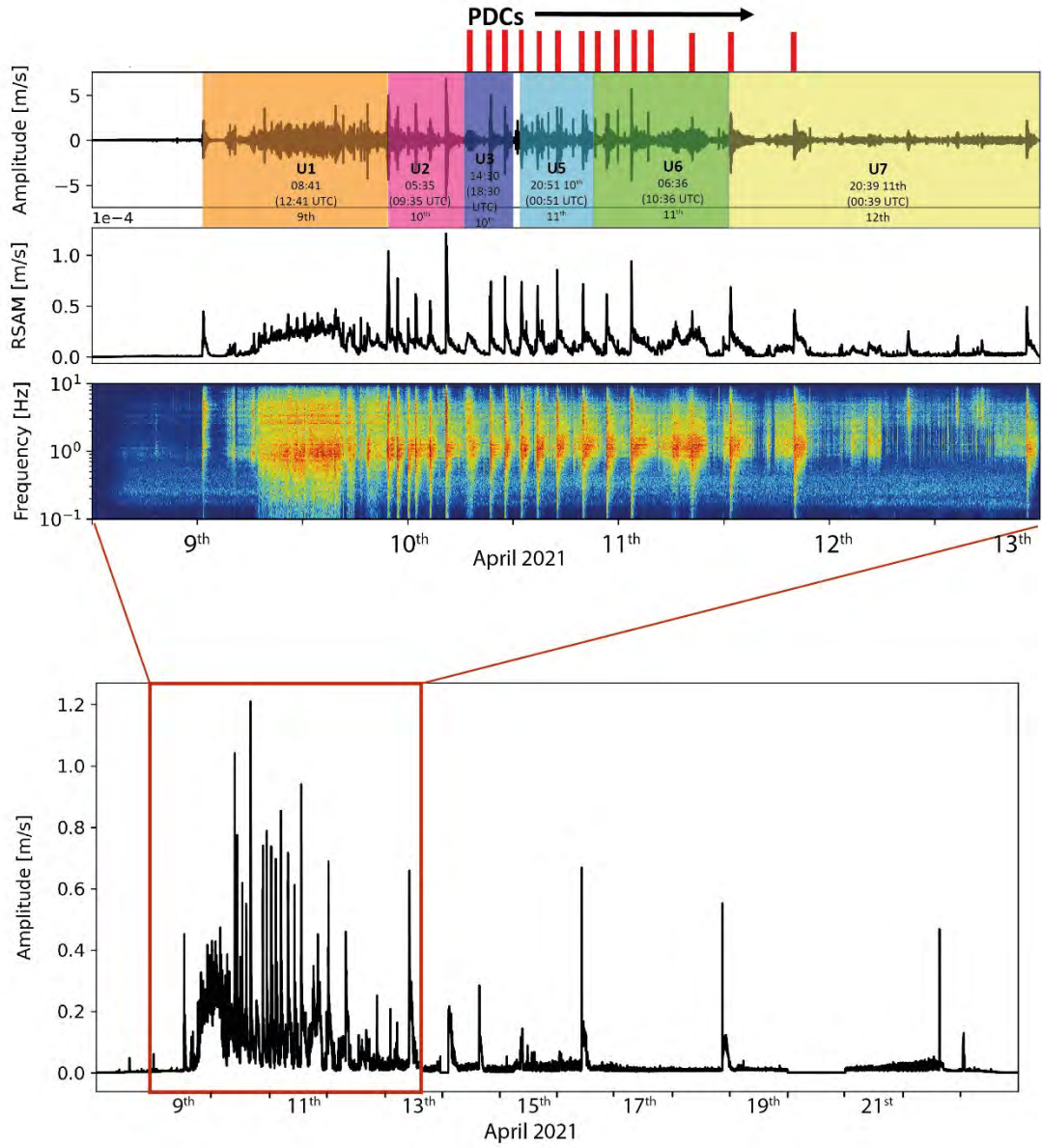
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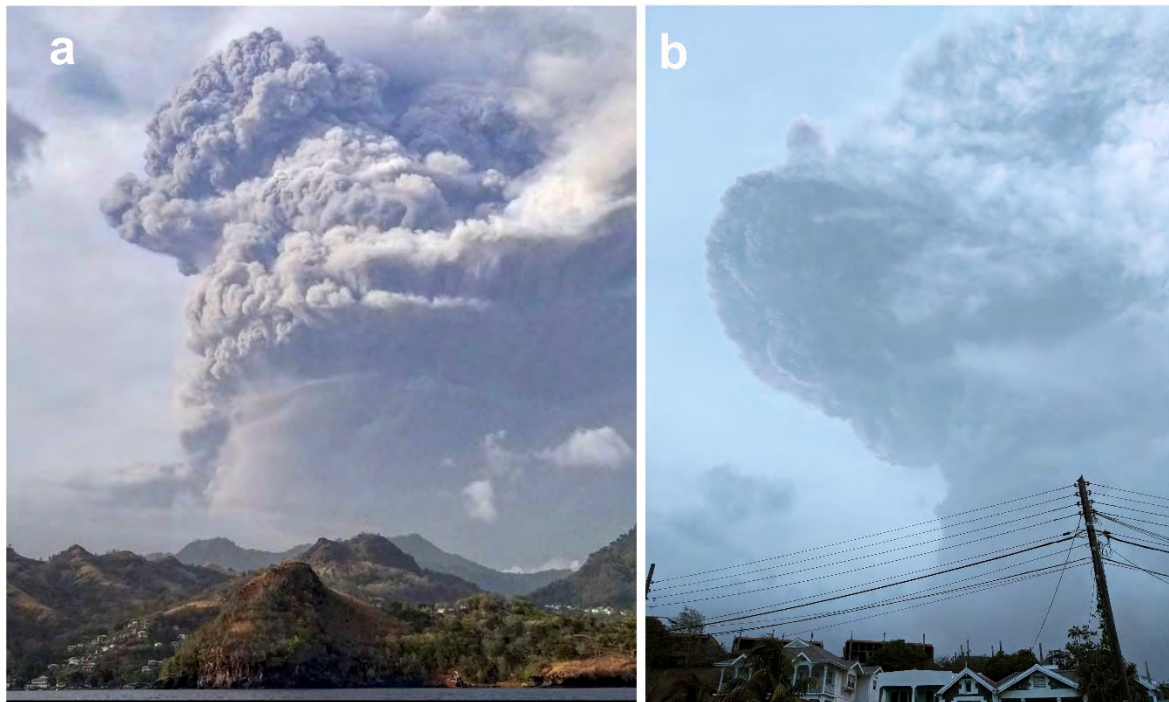
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965 Table 1

Unit	Max thickness (location)	Comments/features	Accretionary lapilli	Components
U1	20 cm (700m SE)	Crudely stratified, moderate GS lapilli (R.G.) Horizons rich in hydrothermally altered clasts	No	S ~10% at base, to 40% at top. Rich in SV clasts, H-rich horizons
U2	20 cm (700m SE)	Ash-rich, up to seven individual layers, Coarser lapilli toward top of unit	No	S ~40-50 wt % clasts
U3	31 cm (600m SE)	Double lapilli, (upper N.G.). Ash layer between lapilli	Yes	S clasts up to 60 wt %)
U4	8 cm (700m SE)	Single ash layer, some diffuse/ scattered lapilli	Yes, abundant	Sparse coarse clasts, mostly D and S.
U5	6 cm (700m SE)	Coarse lapilli, uppermost deposit to East	No	Rich in S, glassy D clasts also prominent
U6	10 cm (2 km SW)	several ash-rich layers	Yes, abundant	
U7	5 cm (4 km SW)	Thin fine grained ash and lapilli	Yes	

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969 **Table 2**

	<b>UNIT 1</b>	<b>UNIT 2</b>	<b>UNIT 5</b>
<b>Total area analysed (cm<sup>2</sup>)</b>	2.19	1.55	1.65
<b>% phenocrysts/microcrysts</b>	24.4	22.6	18
<b>% vesicularity</b>	10	16	32
<b>% vesicularity (crystal free)</b>	<b>27</b>	<b>26</b>	<b>39</b>
<b>No. analysed</b>	32914	22816	10399
<b>Crystal corrected area (mm<sup>2</sup>)</b>	197	131	135
<b>2D BND (per mm<sup>2</sup>)</b>	<b>167</b>	<b>174</b>	<b>77</b>
<b>Average vesicle area (µm<sup>2</sup>)</b>	537	916	4358
<b>Median vesicle area (µm<sup>2</sup>)</b>	<b>109</b>	<b>121</b>	<b>252</b>
<b>Min vesicle area*</b>	3.78	3.76	2.93
<b>Max vesicle area (mm<sup>2</sup>)</b>	0.85	2.80	2.55
<b>Average Elongation</b>	0.3	0.3	0.29
<b>Median Elongation</b>	0.29	0.29	0.28
<b>Pixel size (µm<sup>2</sup>)</b>	0.66	0.86	0.76

970 \*Constrained by minimum pixel size

971 **Table 3**

	<b>Expert 1</b>	<b>Expert 2</b>	<b>Expert 3</b>	<b>Expert 4</b>	<b>Expert 5</b>
<b>Unit 1</b>	<b>15 (12,22)</b>	<b>11 (5,14)</b>	<b>15 (8,20)</b>	<b>18 (14,57)</b>	<b>10</b>
<b>Unit 2</b>	<b>19 (16,29)</b>	<b>15 (12, 16)</b>	<b>16 (12,18)</b>	<b>28 (18, 53)</b>	<b>24</b>
<b>Unit 3</b>	<b>6 (4,12)</b>	<b>5 (4,6)</b>	<b>6 (4,7)</b>	<b>9 (7,16)</b>	-
<b>Unit 4-7</b>	<b>15 (13,18)</b>	<b>12 (11,17)</b>	<b>8 (5,12)</b>		
<b>Total Thickness (From whole tephra thickness isopach)</b>	<b>44 (35,98)</b>	<b>37 (24,82)</b>	<b>31 (29,90)</b>	<b>44 (30,99)</b>	<b>74</b>

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980 Table 4

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	<b>5%ile</b>	<b>Median</b>	<b>95%ile</b>	<b>Mean</b>	<b>+/- StDev</b>
Unit 1	4.6	14.5	51.2	21.6	+/- 15.3
Unit 2	12.2	17.2	47.1	24.2	+/- 12.5
Unit 3	3.8	6.0	15.1	7.9	+/- 3.7
Units 4-7	5.1	12.1	17.7	11.8	+/- 4.0
<b>Summed Units</b>	<b>37.2</b>	<b>63.3</b>	<b>102.4</b>	<b>65.4</b>	<b>+/- 20.4</b>
<b>Total isopach map volume</b>	<b>25.6</b>	<b>42.3</b>	<b>98.0</b>	<b>52.7</b>	<b>+/- 24.3</b>
<b>Mean of Summed and Total</b>				<b>59.1</b>	

992

993 Table 5

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	<b>Volume (x10<sup>6</sup>m<sup>3</sup>) to nearest million</b>
Tephra fallout from isopachs	59 +/- 20
PDC deposits (in valleys outside the Summit Crater)	
Larikai	(6) +/-1.5
Roseau	(4) +/- 1
Upper Wallibou (3 valleys)	(6) +/-1.5
Upper South East flank (upper Rabacca)	(1) +/-0.25
	17 +/-3.5
Intracrater fill (DEM differencing)	43 +/- 10
<b>Total bulk volume</b>	<b>119 +/-24</b>

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1002 Table 6

Event # <sup>(1)</sup>	Day (April 2021)	RSAM Spike Time <sup>(2)</sup> (UTC)	Spike Duration (minutes)	Plume emergence <sup>(3)</sup> (UTC)	Unit	Visual/Stratigraphic observations and correlation with other geophysical datasets
1	9 <sup>th</sup>	12.41	11	12.50	U1	Distinctive explosion and plume, widely observed
2,3,4	9 <sup>th</sup>	18.59		19.00 (start)	U1	Deposit lapilli-rich/ash poor on island with slight coarsening upwards in sequence but individual pulses difficult to differentiate until last 2 explosions shown here.  Following the initial pulse at 19:00 RadiF ash imagery shows near continuous pulsatory plume dispersed ENE, with some distinct pulses towards end of sequence consistent with RSAM spikes at 06:37 and 07:24 UTC.
					U1	
					U1	
	10 <sup>th</sup>			04:30 (End)	U1	
5	10 <sup>th</sup>	06:37	11		U1	
6	10 <sup>th</sup>	07:24	3	07:50	U1	Minor (ash) fallout in SE. island between midnight and 05:30 (09:30 UTC) inferred to be from last two explosions.
7	10 <sup>th</sup>	09.35	23	09.50	U2	Begins with distinct plume seen in SE island at dawn (05:35 LT, 09:35 UTC), with visible fallout. Axisymmetric plume in RadiF. Plume travels ESE.  Minor ashfall begun across Barbados mid-morning, then intensifies, ashfall in Kingstown during day.  Boundary uncertainty: Distinctive ash-rich deposits but with 7 discrete coarse pulses evident in medial locations (six explosions here)
8	10 <sup>th</sup>	10.47	8	11.00	U2	
9	10 <sup>th</sup>	12.02	12	12.10	U2	
10	10 <sup>th</sup>	12.54	13	13.00	U2	
11	10 <sup>th</sup>	14.27	14	14.40	U2	
12	10 <sup>th</sup>	16.20	23	16.30	U2	
13	10 <sup>th</sup>	18.30	13	18.50	U3	
14	10 <sup>th</sup>	21.20	19	21.30	U3	

15	10 <sup>th</sup>	23.02	20	23.10	U4	Ash-rich layer with some coarser lapilli. Although U4 could be predominantly from PDC fallout, distributed across island.  Boundary uncertainty: lower peak energy (RSAM) explosions could implying PDC or lower intensity ash generation
16	11 <sup>th</sup>	00.51	21	01.00	U5	Coarse scoria-rich deposits. Prominent in deposits across island, forming carapace on Windward sequences.  3 <sup>rd</sup> Explosion in this sequence cuts power across the island.  Early morning rain in Barbados, combined with continued ash fall.  Boundary Uncertainty: nature of U5 deposit (large scoria lodged in finer ash) makes precise no. of explosions harder to determine. Daylight visual observations more consistent with U6 explosions.
17	11 <sup>th</sup>	02.44	16	02.50	U5	
18	11 <sup>th</sup>	04.59	11	05.10	U5	
19	11 <sup>th</sup>	07.55	13	08.10	?U5	
20	11 <sup>th</sup>	10.36	13	10.50	U6	Number of fine-grained ash layers, rich in accretionary lapilli. Limited to Western flanks of volcano. PDC generation continues at this time.  Eruption columns predominantly carried to Leeward in lower atmospheric winds from this time (below windshear). Rainfall on Windward side of island (lahars) –little further ash deposition on this side of island & ‘clearing air’ during morning. Steaming in Rabacca reported at this point.
21	11 <sup>th</sup>	13.24	22	13.40	U6	
22	11 <sup>th</sup>	18.11	11	18.20	U6	
23	11 <sup>th</sup>	20.22	12	20.20	U6	
24	12 <sup>th</sup>	00.39	18	00.50	U7?	Majority of this sequence is manifest in fine indistinct ash sequence of U7 (when present at all – only found proximal to volcano).  N. Leeward observers report no ash by 13 <sup>th</sup> April. PDCs still observed.  Rainfall continues on 12 <sup>th</sup> and 13 <sup>th</sup> April.  Ashfall in Barbados now negligible.  We infer that there would be almost no fallout visible in our sequence from the last three explosions.
25	12 <sup>th</sup>	07.59	20	08.10	U7 ?	
26	12 <sup>th</sup>	20.53	24	21.30	U7?	
27	13 <sup>th</sup>	10.23	24	10.40	U7?	
28	14 <sup>th</sup>	02.27	17	02.40	U7 ?	
29	14 <sup>th</sup>	15.31	10	15.40	U7?	
30	16 <sup>th</sup>	10.16	5	10.50	-	
31	18 <sup>th</sup>	20.49	10	21.00	-	
32	22 <sup>nd</sup>	15.09	9		-	

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