

2021-05-01

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<http://hdl.handle.net/10026.1/17542>

10.1088/1367-2630/abfab2

New Journal of Physics

IOP Publishing

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Exact solutions in radiation reaction and the radiation-free direction

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We present new exact solutions of the Landau-Lifshitz and higher-order Landau-Lifshitz equations describing particle motion, with radiation reaction, in intense electromagnetic fields. Through these solutions and others we compare the phenomenological predictions of different equations in the context of the conjectured ‘radiation-free direction’ (RFD). We confirm analytically in several cases that particle orbits predicted by the Landau-Lifshitz equation indeed approach the RFD at extreme intensities, and give time-resolved signals of this behaviour in radiation spectra.

I. INTRODUCTION

Despite having been studied for more than a century [1–3], radiation reaction (RR) continues to attract theoretical [4–9], computational [10, 11], and experimental [12–15] interest. In large part, this attention is driven by intense laser systems [16–19] now granting access to regimes where quantum and classical RR forces can dominate the Lorentz force. For recent reviews see Refs. [20, 21].

It is well-known that including RR effects allows for new phenomena, such as anomalous particle trapping [22], chaotic motion [23], symmetry breaking [24, 25], and significantly enhanced generation of certain plasma wave modes [26]. In such scenarios it is often expected that classical RR effects receive significant quantum corrections; the simpler setting of classical physics can, nevertheless, still provide important insight [27], and classical effects can persist in quantum theory [22]. From a theoretical point of view, classical radiation reaction also remains an interesting test-bed for the emergence of non-perturbative physics [3, 7, 27, 28].

Eliminating the radiation fields, created by the charges, from the classical equations of motion, one arrives at the Lorentz-Abraham-Dirac (LAD) equation [1–3] which contains the third time derivative of position. This implies unwanted effects such as runaway solutions and pre-acceleration. The Landau-Lifshitz [29] (LL) equation, obtained from LAD through ‘reduction of order’ (see below) is however free of these difficulties and is typically an excellent approximation below scales where quantum effects appear [5].

Given the subtlety of unphysical non-perturbative effects, and the number of complex phenomena attributable to RR, exact solutions can help in making precise statements. In this paper we present two new exact solutions for field configurations that depend only on a single lightlike direction, with polarisation either longitudinal or transverse to that direction. Specifically, for the longitudinal case we give the exact solution of the LL equation; and for transverse polarisation, i.e. plane waves, we solve

the *second-order* Landau-Lifshitz equation obtained by iteration of reduction of order in closed form, in a physically motivated setup.

As an application of these exact solutions (and others to be discussed), we provide analytic evidence confirming the tendency of radiation reaction in strong fields to align particle motion with the ‘radiation free direction’ (RFD) in which they locally experience zero acceleration transverse to their direction of motion, thus minimising radiation losses. The RFD hypothesis was suggested and supported by numerical simulations in [30].

This paper is organised as follows. We first give notation and conventions, and write down the Lorentz force, LAD and LL equations for reference. In Sect. II we solve the LL equation in longitudinally polarised electric fields of arbitrary strength and form, and demonstrate analytically that orbits transition to the RFD. We also show that RFD dynamics distinguishes between other proposed classical equations of motion. In Sect. III we investigate RFD dynamics in the case of plane wave, i.e. transversely polarised backgrounds, for which the solution of the LL equation is already known, and look for signals of (the transition to) RFD dynamics in emission spectra. In Sect. IV we solve the ‘second-order’ LL-like equation, obtained from iteration of reduction of order, in plane waves and compare its predictions with those of the standard LL equation. We conclude in Sect. V.

A. Notation and conventions

We use Heaviside-Lorentz units with $c = \hbar = 1$ and employ lightfront coordinates $x^\pm := x^0 \pm x^3$ and $x^\perp := (x^1, x^2)$. Lightfront momentum components p^\pm and p^\perp are defined analogously. It is convenient to introduce a lightlike vector n^μ such that $x^+ = n \cdot x$.

In a background field $F_{\mu\nu}$ the LAD equation of particle motion is obtained from the coupled system of the Lorentz force law and Maxwell’s equations by integrating out the dynamical electromagnetic fields. Writing $f \equiv eF/m$, the LAD equation is

$$\ddot{x}^\mu = f^{\mu\nu} \dot{x}_\nu + \tau_0 \mathcal{P}^{\mu\nu} \ddot{x}_\nu, \quad (1)$$

where a dot is a proper-time derivative, $\tau_0 := e^2/6\pi m$ is the characteristic timescale for RR, and $\mathcal{P}^{\mu\nu}$ projects orthogonally to \dot{x}^μ . The standard Lorentz force equation

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for motion in the background only, i.e. neglecting radiation and RR, is recovered by setting $\tau_0 = 0$.

The LL equation is obtained by substituting (1) back into itself to eliminate the third derivative in favour of new, explicitly f -dependent terms, and then neglecting terms of order τ_0^2 and larger. The LL equation thus found is

$$\ddot{x}^\mu = f^{\mu\nu}\dot{x}_\nu + \tau_0 \dot{f}^{\mu\nu}\dot{x}_\nu + \tau_0 \mathcal{P}^{\mu\nu} f_{\nu\rho} f^{\rho\sigma} \dot{x}_\sigma. \quad (2)$$

This process is called ‘reduction of order’, since it replaces a third-order ODE, the LAD equation with a second-order ODE. The process can of course be extended to higher orders in τ_0 ; we return to this in Sect. IV.

II. LONGITUDINAL POLARISATION

We consider first electric fields depending only on x^+ , which represent electromagnetic pulses propagating in the negative z direction. We take the fields to be ‘longitudinally’ polarised in the z -direction [31, 32]. This is not a solution of the source-free Maxwell equations, but waves of this type can be realised in a plasma or using binary optics for light [33]. The non-zero components of the field tensor are

$$F_{+-} = -F_{-+} = \partial_+ A_-(x^+), \quad (3)$$

equivalent to having a purely electric field $\mathbf{E} = (F_{-+}/2)\hat{\mathbf{z}}$. This class of fields includes of course constant electric, on which we comment below. Recall that any massive particle orbit can be parameterised by lightfront time [34]. Using this, the solution of the Lorentz force equation in our background is easily found; writing $a(x^+) = eA^+$, the first integrals, i.e. the particle momenta $\pi^\mu = m\dot{x}^\mu$, are

$$\text{(Lorentz)} \quad \pi^+(x^+) = p^+ - a(x^+), \quad (4)$$

$$\text{(Lorentz)} \quad \pi^\perp(x^+) = p^\perp, \quad (5)$$

with initial conditions $\pi^\mu = p^\mu$ at some initial time $x^+ = x_0^+$ before the field turns on. π^- follows from the mass-shell condition, $\pi^+\pi^- - \pi^\perp\pi^\perp = m^2$, and the momenta determine the orbit, $x^\mu(\tau)$, via quadratures.

Observe that we have uniform motion perpendicular to the pulse direction in (5), for any p^\perp , as the ‘transverse’ directions x^\perp decouple. According to [30], though, when RR is included the particle should move toward the RFD which, for the field (3), is parallel or anti-parallel to the electric field polarisation according to the sign of the charge, in other words the negative or positive z direction. To see the impact of RR effects, we turn to the LL equation.

We consider first the transverse components of (2):

$$\begin{aligned} \ddot{x}^\perp &= -\tau_0 \dot{x}^\perp \dot{x}^- (2f_{+-})^2 \dot{x}^\perp \\ &= -\tau_0 (1 + \dot{x}^\perp \dot{x}^\perp) (2f_{+-})^2 \dot{x}^\perp, \end{aligned} \quad (6)$$

using the mass-shell condition in the second line. It is clear from (6) that $\ddot{x}^\perp < 0$, meaning that that RR effects drive

the transverse velocity \dot{x}^\perp monotonically to zero. As a result, the motion becomes confined to the tz plane, which confirms that the particle momentum indeed becomes aligned with the RFD. A complementary argument is to note that the transverse equation of motion (6) always has the trivial solution $\dot{x}^\perp = 0$; a stability analysis then shows that this solution represents an ‘attractor’, meaning any deviation from $\dot{x}^\perp = 0$ will be killed by virtue of (6).

Notably, we find that the system on the attractive sub-manifold, $\dot{x}^\perp = 0$, is integrable; \dot{x}^- becomes trivially determined from the mass-shell condition, $\dot{x}^- = 1/\dot{x}^+$, and (2) reduces to a single equation for \dot{x}^+ ,

$$m\ddot{x}^+(x^+) = -\dot{x}^+ \partial_+ a - \tau_0 \partial_+^2 a, \quad (7)$$

which can be solved exactly. Changing independent variable from proper time to lightfront time, (7) becomes

$$\partial_+ \pi^+ = -\partial_+ a(x^+) - \frac{\tau_0}{m} \pi^+ \partial_+^2 a(x^+), \quad (8)$$

which is readily solved to yield, for $\pi^+(-\infty) = p^+$,

$$\pi^+(x^+) = e^{-\frac{\tau_0 a'(x^+)}{m}} \left(p^+ - \int_{-\infty}^{x^+} e^{\frac{\tau_0}{m} a'(y)} a'(y) dy \right). \quad (9)$$

As a check, we note that for a constant electric field, $a'' = 0$, and $p^+ = 0$, the LL solution (9) and the Lorentz solution (4) agree, and also solve the LAD equation. This is consistent with the literature result that there is ‘no radiation reaction for hyperbolic motion’ [35–37]; there is though radiation [35, 38, p. 399].

With (9) one can make explicit the stability of the RFD solution. Linearising the LL equation in π^\perp , the equations for π^\pm are unchanged, (since π^\perp enters them *quadratically*), which allows us to solve (6) as

$$\pi^\perp(x^+) \simeq \pi^\perp(x_0^+) \exp \left[-4m\tau_0 \int_{x_0^+}^{x^+} f_{+-}^2 / \pi^+ dy \right], \quad (10)$$

where π^+ is as in (9); thus deviations of motion from the RFD are exponentially suppressed.

A. Example: Sauter pulse

To understand the physics of the exact solution (9) in non-constant fields we turn to an explicit example with a chosen field configuration. We consider a Sauter pulse of width $1/\omega$ and peak field strength determined by a_0 ,

$$a'(x^+) = \frac{m\omega a_0}{2} \text{sech}^2(\omega x^+) \quad (11)$$

$$a(x^+) = \frac{ma_0}{2} (1 + \tanh(\omega x^+)), \quad (12)$$

with positive/negative values of a_0 corresponding to an electric *force* on the particle parallel/anti-parallel to the

direction of pulse propagation (which, recall, is the negative z -direction in our conventions).

$$\pi^+(x^+) = e^{-\frac{\omega\tau_0 a_0}{2} \operatorname{sech}^2(\omega x^+)} \left[p^+ - m \sqrt{\frac{\pi a_0}{2\omega\tau_0}} e^{\frac{\omega\tau_0 a_0}{2}} \left(\operatorname{erf} \sqrt{\frac{\omega\tau_0 a_0}{2}} \tanh \omega x^+ + \operatorname{erf} \sqrt{\frac{\omega\tau_0 a_0}{2}} \right) \right], \quad (13)$$

while for $a_0 < 0$, absolute value signs should be inserted under the square roots and erf should be replaced by erfi. The corresponding Lorentz force results are given simply by inserting (12) into (4).

Note that the solutions of both the LL and Lorentz equations are valid only as long as $\pi^+ > 0$, as is required for massive particles. For $\pi^+ \rightarrow 0$, the particle is accelerated to almost co-propagate with the field, and reaches the speed of light in finite *lightfront* time, after which the particle ‘leaves the spacetime manifold’ [31, 32]. Note that this finite lightfront time corresponds to *infinite* lab-frame, or proper, time. As such, considering both (4) and (9), it is clear that the sign of a_0 is a crucial factor in determining properties of the motion.

We consider first the case $a_0 < 0$, for which the force on the particle is anti-parallel to the propagation direction; for a head-on collision, this means the particle is accelerated in its direction of initial propagation. We define the momentum transfer W by

$$W := \pi^+(\infty) - \pi^+(-\infty) = \pi^+(\infty) - p^+. \quad (14)$$

Without RR, $W_{\text{Lorentz}} = m|a_0|$, but according to the LL equation RR reduces the momentum transfer; one finds from (13) that

$$W_{\text{LL}} \simeq \frac{m^2}{\alpha\omega} \quad \text{for} \quad \frac{\alpha\omega|a_0|}{m} \gg 1. \quad (15)$$

The behaviour of W as a function of a_0 , for both the Lorentz and LL equations of motion, is shown in Fig. 1.

We turn to the case $a_0 > 0$, for which the force on the particle is parallel to the field propagation direction (and therefore opposite the initial propagation direction of the particle). In this case we find that for sufficiently large a_0 , the particle can be brought to rest, and for even larger a_0 can be caught in the pulse and accelerated almost to the speed of light. In the Lorentz case, the respective thresholds above which these phenomena occur are easily read off from (4) as

$$a_{0,\text{stop}} = p^+/m - 1 \quad \text{and} \quad a_{0,c} = p^+/m. \quad (16)$$

Including RR through the LL equation, the thresholds are the solutions of an intractable transcendental equation, but it is easily found by numerical investigation that both thresholds are lowered by RR. What this means is that there is a range of a_0 such that particles are back-scattered

For this field the integral in (9) can be performed analytically in terms of the error function with the result, for $a_0 > 0$,

by RR effects, i.e. their direction of motion is reversed. Example orbits illustrating this effect are plotted in Fig. 2.

In the figure, an unphysically large value of α has been used to exaggerate the effect of RR (along with an ω much larger than is phenomenologically motivated). This is because the solution (9) describes a particle already moving in the RFD, and because longitudinal forces produce much less radiation than transverse forces, RR is consequently a very small effect. RR will naturally be more pronounced in the transition to the RFD, i.e. as π^+ is driven to zero and, when close to zero, follows (10). Rather than analyse the radiation spectra from this approximate final stage motion, though, we will consider in the next section the case of plane waves, which admit exact solutions to the LL equation for arbitrary initial conditions. Before doing so we compare the results above with those of other classical equations.

B. Alternative equations

We ask here what different classical equations predict regarding transition to, and motion in, the RFD. The question is somewhat academic, as unlike LAD (the classical equation of motion) or LL (an approximation to it)

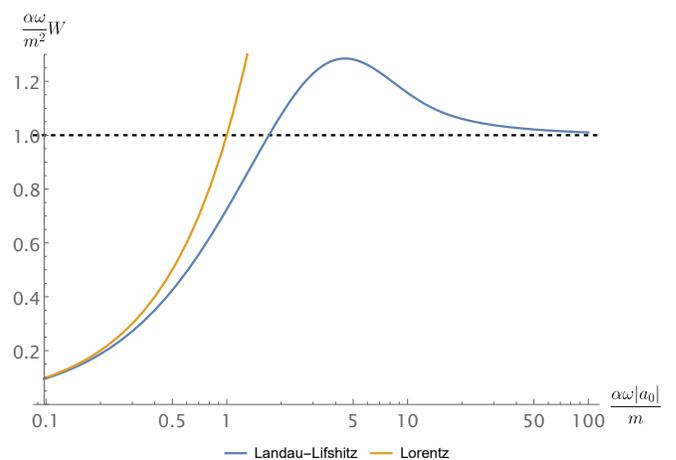


FIG. 1. Momentum transfer W (14), in a head-on collision with a Sauter pulse with $a_0 < 0$. Without RR, the momentum transfer is unbounded. With RR, W approaches a constant at large $|a_0|$, as RR effects dominate over the electric force.

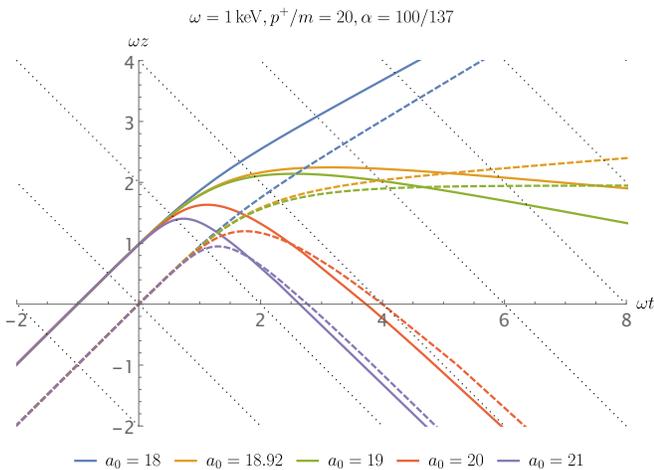


FIG. 2. Particle orbits in a head-on collision with a longitudinally polarised Sauter pulse. Solid: LL, dashed: Lorentz. The LL orbits have been displaced by unity in the ωx^- direction to distinguish them from the Lorentz orbits. Note that an unphysically large α has been used to exaggerate RR.

other equations proposed over the years bring in some external assumption. We will therefore be brief.

Eliezer's equation [39] (rederived by Ford-O'Connell in a slightly different context [40]) is

$$\ddot{x}^\mu = f^{\mu\nu} \dot{x}_\nu + \tau_0 \mathcal{P}^{\mu\nu} \frac{d}{d\tau} (f_{\nu\rho} \dot{x}^\rho). \quad (17)$$

Linearising in the transverse momentum as above, we find

$$\dot{\pi}^\perp \simeq \left(\frac{\tau_0}{m} \frac{\dot{\pi}^+}{\pi^+} \partial_+ a \right) \pi^\perp, \quad \pi^+ = \pi_{\text{LL}}^+, \quad (18)$$

This means that motion *in* the RFD, when it is reached, is exactly as described by LL. The approach to the RFD is, however, not monotonic: direct numerical investigation confirms that the coefficient of π^\perp in (18) is initially negative, but can change sign, implying attraction to or repulsion from the RFD.

Ultimately, it is impossible to turn off quantum effects, and the predictions of some proposed equations do not agree with the classical limit of QED results [41–43]. An example is the Mo-Papas equation [44],

$$\ddot{x}^\mu = f^{\mu\nu} \dot{x}_\nu + \tau_0 \mathcal{P}^{\mu\nu} f_{\nu\rho} \ddot{x}^\rho. \quad (19)$$

Linearising again, one finds in this case that

$$\dot{\pi}^\perp \simeq -\frac{\tau_0}{m^2} (\partial_+ a)^2 \pi^\perp, \quad \pi^+ = \pi_{\text{Lorentz}}^+ \quad (20)$$

The first of these equations tells us that, for small transverse momentum, $\dot{\pi}^\perp = 0$ is again a stable attractor, and the orbit goes to the RFD. However, motion in the RFD is that predicted by the Lorentz force equation, without radiation reaction. Thus the LL, Eliezer and Mo-Papas, equations all predict different behaviour, and it seems that, in principle, RFD dynamics can differentiate between classical equations of motion.

III. TRANSVERSE POLARISATION

We turn now to plane waves, i.e. functions of $x^+ = n \cdot x$, which are transversely, rather than longitudinally polarised. Their potential $a_\mu := eA_\mu$ can always be written $a_\mu = m a_0 \delta_\mu^\perp f_\perp(\omega x^+)$ in which ω is some frequency scale, f_\perp describes the shape of the field, and the dimensionless invariant a_0 characterises the peak field strength.

The solution to the LL equation was first given for monochromatic, linearly polarised plane waves in Ref. [45], and extended to all plane waves in Ref. [46]. It is conveniently presented using the following parameterisation of the momentum π_μ ,

$$\pi^\mu = \frac{1}{h} \left(p^\mu - B^\mu + n^\mu \frac{2p \cdot B - B^2 + m^2(h^2 - 1)}{2n \cdot p} \right), \quad (21)$$

in which B_μ is purely transverse. Writing $\phi \equiv \omega x^+$, $k_\mu \equiv \omega n_\mu$ and a prime for a ϕ -derivative, the functions h and B_μ are

$$h = 1 + \tau_0 \frac{k \cdot p}{m^4} \int_{-\infty}^{\phi} dy a'(y) \cdot a'(y), \quad (22)$$

$$B_\mu = \tau_0 \frac{k \cdot p}{m} a'_\mu + \int_{-\infty}^{\phi} dy h(y) a'_\mu(y), \quad (23)$$

with initial conditions $\pi^\mu(-\infty) = p^\mu$. The parameterisation (21) is chosen because it makes clear that *all* momentum components are actually proportional to $1/h = \pi^+/p^+$; in this sense π^+ governs the dynamics.

It is more convenient to analyse the transition to the RFD at the level of the solution (21), rather than at the level of the LL equation, as we did for longitudinally polarised fields. Using (22)–(23), we see that the ratio of transverse momentum to z -momentum,

$$\frac{|\pi^\perp|}{\pi^z} = \frac{2|\pi^\perp|}{\pi^+ - \pi^-}, \quad (24)$$

behaves as $\sim a_0^{-3}$ for $a_0 \gg 1$. Hence, while motion in a plane wave can never be confined exactly to the tz plane due to the explicitly transverse polarisation, motion is *dominantly* in the z direction for $a_0 \gg 1$. This is indeed the radiation-free direction according to Ref. [30],

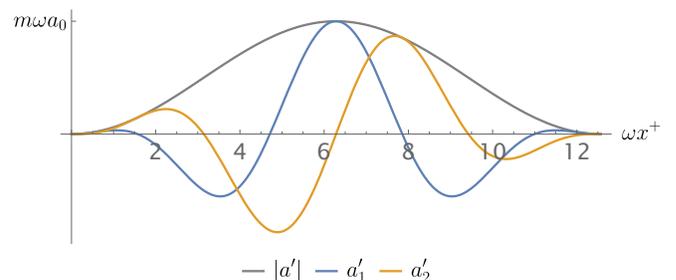


FIG. 3. Electric field of a circularly polarised pulse with a \sin^2 envelope, cf. (25).

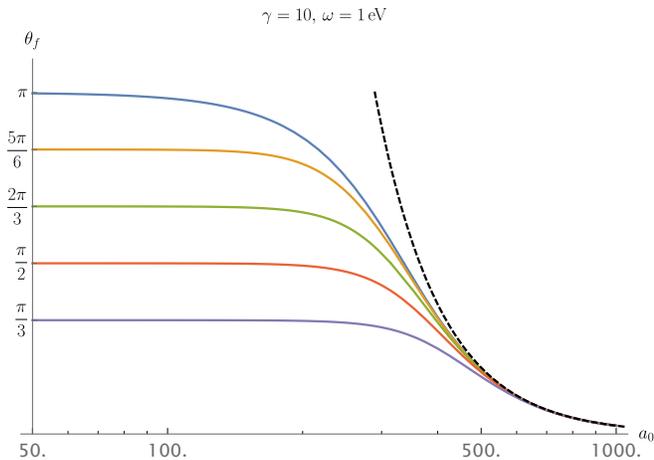


FIG. 4. Outgoing scattering angle, relative to the laser propagation direction, as a function of a_0 , at fixed initial particle energy and for several incoming collision angles, indicated at each solid curve, in the pulse (25). The asymptotic $1/a_0^3$ scaling is also shown (dashed). (The scattering angle is essentially independent of a_0 below the range displayed.) Optical frequency, $\omega = 1$ eV, is chosen as being the most phenomenologically relevant; the asymptotic behaviour then sets in for an a_0 of several hundred, owing to the smallness of $\omega\tau_0 \simeq 10^{-8}$.

although the limits of the formulae provided there need to be taken with care as, in a plane wave, both field invariants vanish.

We illustrate the particle motion using a circularly polarised, few-cycle pulse with a \sin^2 envelope, for which the integrals in (22) and (23) can be performed analytically. The pulse is given by

$$\begin{pmatrix} a'_1 \\ a'_2 \end{pmatrix} = ma_0 \sin^2 \frac{\phi}{4} \begin{pmatrix} \cos \phi \\ \sin \phi \end{pmatrix} \quad (25)$$

for $0 \leq \phi \leq 4\pi$, and vanishing otherwise, see Fig. 3; the envelope ensures that the electromagnetic fields, and their derivatives, vanish at the edge of the pulse.

Performing the integrals in (22) and (23), we can finesse the approximation (24) and calculate the outgoing scattering angle θ_f of the incident particle with respect to the laser propagation direction. We find

$$\theta_f \approx \frac{24}{7\pi\alpha a_0^3 \omega} \quad a_0 \gg 1, \quad (26)$$

which, note, is completely independent of the *incoming* direction of the particle, and is clearly in agreement with (24). Fig. 4 shows the full a_0 -dependence, and the asymptote (26), for several incidence angles.

This confirms that, at high intensity, all particles are driven to propagate dominantly in the RFD (with relatively small transverse momentum). We illustrate the transition to (near) laser-collinear scattering by plotting particle orbits in the (t, z) plane in Fig. 5. Note that the parameters in Fig. 5 have been chosen for visual clarity rather than physicality: for realistic parameters and a_0

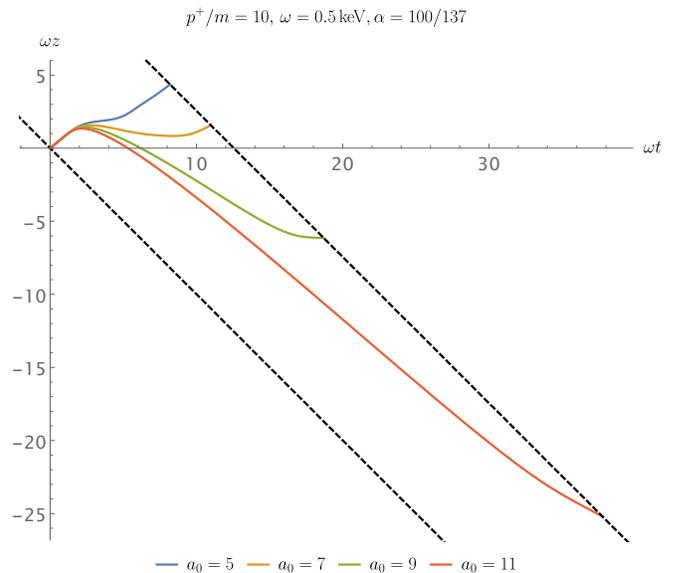


FIG. 5. Particle orbits for a head-on collision with the pulse (25), whose spacetime extent is indicated by the dashed lines. As a_0 increases, the particle goes from scattering forward, to coming to a stop, to scattering backward. (Note: an unphysically large α has been used to exaggerate RR.)

in the radiation-dominated regime, the particle is carried with the pulse for many cycles' worth of *lab* time.

A. Radiation spectra

Signatures of RFD dynamics can be found in the emitted radiation spectra. The radiated photon 4-momentum K^μ has the standard form

$$K^\mu = - \int \frac{d^2\ell^+ d\ell^+}{(2\pi)^3 2\ell^+} \ell^\mu |j(\ell)|^2, \quad (27)$$

where j_μ is the current in Fourier space,

$$j_\mu(\ell) = -e \int dx^+ e^{i\ell \cdot X(x^+)} \frac{\partial}{\partial x^+} \left(\frac{\pi_\mu}{i\ell \cdot \pi} \right), \quad (28)$$

and X^μ is the particle orbit. All but one of the integrals in (27) can be performed exactly; the final expression for K^μ is

$$K^\mu = \frac{2}{3} \frac{e^2}{4\pi} \frac{k \cdot p}{m^2} \int d\phi \frac{m^2 h'(x)^2 - B'(x)^2}{m^4 h(x)^3} \pi^\mu(\phi). \quad (29)$$

Examining the integrand shows us from where in the pulse radiation is generated. For low a_0 RR is negligible, and we find that emission of all spectral components K^μ is supported mainly near the peak of the pulse. A signature of the transition to the RFD is then that emission into different spectral components K^μ becomes *time-resolved* at high intensity; as shown in Fig. 6(a), radiation emitted anti-collinear/transverse/collinear to the laser is predominantly emitted near the beginning/peak/end of the pulse.

This corresponds directly to the change in particle direction associated with transition to the RFD, as sketched in Fig. 6(b).

In Fig. 6(a) the components are normalised to their maxima to highlight their different temporal supports. To extract analytic estimates for their relative sizes, when $a_0 \gg 1$, requires a little care. The components K^+ , K^- scale according to naive estimates using (21), viz.,

$$K^\perp = \mathcal{O}(a_0) \quad \text{and} \quad K^- = \mathcal{O}(a_0^4). \quad (30)$$

However, the dominant contribution to K^+ does not scale as a_0^{-2} as might be expected; the dominant contribution comes from a boundary term hidden in the contribution from B'^2 and is [27]

$$\frac{K^+}{p^+} \sim -\frac{2}{3} \frac{e^2}{4\pi} \frac{k \cdot p}{m^4} \int d\phi \frac{h^2 a'^2}{h^4} = -\int d\phi \frac{h'}{h^2} = 1. \quad (31)$$

This scaling behaviour is a consequence of conservation of total lightfront momentum; K^+ and π^+ both being non-negative, the radiation field cannot carry off more than the initial p^+ of the particle.

While the total perpendicular momentum is also conserved, π^+ is not constrained and can always absorb an arbitrarily large recoil. Finally, due to the lack of symmetry in x^+ the collinear momentum is simply not conserved, putting no restriction on the particle ‘‘pumping’’ laser momentum into its own radiation field.

It may seem paradoxical that, comparing (31) and (30), the dominant component is in the RFD, but we must remember that motion in the RFD minimises radiation for a given magnitude of the applied force [30]. In the radiation-dominated regime, though, the particle sheds all of its lightfront momentum early in the rise of the pulse. As the particle approaches the radiation-free direction the magnitude of the force and, eventually, the Lorentz factor increase, both of which strongly enhance the radiated power.

IV. LL TO SECOND ORDER

The LL equation (2) has been obtained by iterating the LAD equation (1) to *first* order in τ_0 . To assess the quality of this approximation we consider the size and influence of corrections by performing the iteration to second order in τ_0 . We will refer to the LAD equation iterated to n :th order in τ_0 as LL_n (so that the Lorentz equation of motion is LL_0). A somewhat lengthy calculation yields the following expression for LL_2 ,

$$\begin{aligned} \ddot{x}^\mu = & (f + \tau_0 \dot{f} + \tau_0^2 \ddot{f} - 2\tau_0^2 (\dot{x} f^2 \dot{x}) f)^{\mu\nu} \dot{x}_\nu \\ & + \tau_0 \mathcal{P}^{\mu\nu} [f_{\nu\rho}^2 + 2\tau_0 f_{\nu\rho}^3 + 2\tau_0 (f^2)_{\nu\rho}^*] \dot{x}^\rho. \end{aligned} \quad (32)$$

A. Transverse polarisation

As for LL_1 , if one can solve LL_2 for π^+ in a plane wave, as a function of x^+ , then the transverse momenta π^\perp

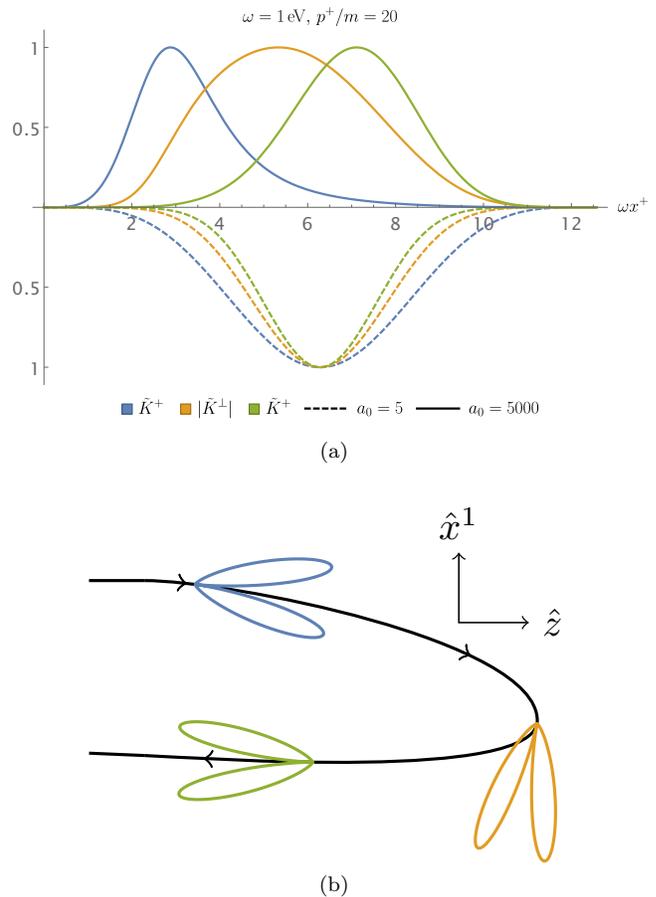


FIG. 6. Radiation reaction makes the radiation profile in a pulse time-resolved. (a): The radiated 4-momentum in the pulse (25), as a function of lightfront time. Each component has been normalised to its maximum, $\tilde{K}_\mu = K_\mu / \max |K_\mu|$, and the dashed curves have been reflected in the horizontal axis for clarity. (b): Sketch of the particle orbit and the radiation cone rotating along it.

are easily calculated, and π^- is given by the mass-shell condition. We therefore focus on π^+ .

The momentum π^μ is again conveniently parameterised as in (21), so that we again have $\pi^+/p^+ = 1/h$. It is useful to define the two dimensionless RR parameters

$$\delta = \frac{2}{3} \frac{e^2}{4\pi} \frac{k \cdot p}{m^2}, \quad \Delta := a_0^2 \delta, \quad (33)$$

where δ is essentially an energy parameter, which should strictly be small in the classical regime, whereas Δ depends on the field strength, and so can be large. Denoting ϕ -derivatives with a prime as before, the LL_2 equation for h becomes

$$h' = -\frac{\delta}{m^2} a' \cdot a' - \frac{4\delta^2}{m^2 h} a' \cdot a'', \quad (34)$$

which is nonlinear due to the factor of h^{-1} in the final term: this is an Abel equation of the second kind, and thus not analytically solvable in general, while special cases are usually only solvable parametrically [47].

However, there is a solvable case which is relevant to the physical situation of interest. We have seen that at high intensities, a particle can be stopped and turned around soon after it enters the pulse, i.e. before reaching the peak. We therefore consider dynamics in the initial *rise* of the pulse, a simple model of which is

$$a' \cdot a' = -m^2 a_0^2 e^\phi, \quad \phi < 0. \quad (35)$$

The individual components of a' may contain oscillatory factors, c.f. the brackets in (25), without affecting (35). (Simple extensions are to continue directly to $\phi > 0$ as a model of an ever-increasing field, or use $e^{-|\phi|}$ to model a sharply peaked field [48, 49].) With this choice, we bring (34) to standard Abel form by changing the independent variable from ϕ to

$$z := 1 - \delta \int_{-\infty}^{\phi} dy a'(y) \cdot a'(y) = 1 + \Delta e^\phi, \quad (36)$$

which, note, is just the solution for h from LL₁. Our ODE (34) reduces to

$$h(z) \partial_z h(z) - h(z) = 2\delta. \quad (37)$$

Setting all *explicit* factors of $\delta \rightarrow 0$ in this expression recovers the LL₁ solution; this implies, as suggested in Ref. [50] that the difference between LL₁ and LL₂ depends essentially on δ (which should be small in the classical regime), rather than the potentially large Δ . To confirm this, we need the solution of (37). This is easily found by first solving the equation for z as a function of h . Choosing initial condition $h(z=1) = 1$, the solution is

$$z = h - 2\delta \log \frac{h + 2\delta}{1 + 2\delta}. \quad (38)$$

Again, setting $\delta = 0$ on the right correctly recovers the LL₁ result: the difference is *explicitly* dependent only on δ . Rearranging, we can write h in terms of the Lambert function [51], or product logarithm, W ,

$$-\frac{h(\phi)}{2\delta} = 1 + W \left(-\frac{1 + 2\delta}{2\delta} \exp \left[-1 - \frac{1 + \Delta e^\phi}{2\delta} \right] \right). \quad (39)$$

Interestingly, the argument of W is in the range $(-1/e, 0)$ which means we must take $W \equiv W_{-1}$, i.e. *not* the principal branch of W . As is clear from the expressions above, and as shown explicitly in Fig. 7, the difference between h as predicted by LL₂ vs. LL₁ is extremely small unless the energy parameter δ is taken to be very large, and therefore outside the classical regime.

B. Longitudinal polarisation

For the longitudinally polarised field, (32) implies the following LL₂ equation for the transverse velocity components,

$$\dot{u}^\pm = -4\tau_0 u^+ u^- \left[(1 + 2\tau_0 u^+ \partial_+) (f_{+-})^2 \right] u^\pm. \quad (40)$$

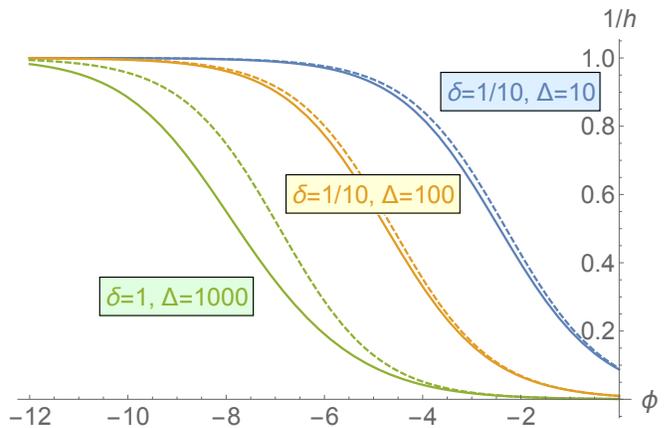


FIG. 7. Momentum $\pi^+/p^+ \equiv 1/h$ in the rise of the pulse, for various δ and Δ , in LL₂ (solid lines) and LL₁ (dashed lines). Increasing the field strength, so Δ , leads to significant momentum loss earlier in the pulse. The rate of this loss is corrected by effects of $\mathcal{O}(\delta)$ in LL₂.

Comparing to (6), the difference is the derivative term inside the brackets; this does not have a definite sign, and so the approach to the RFD is no longer uniform. However, this new NLO term is roughly of order $\alpha\omega p^+/m^2$ relative to the LO term, so is again dependent on energy rather than field strength. As such, higher-order corrections are only significant when the energy parameter is large. We conclude therefore that the RFD hypothesis continues to hold (at least) for parameters where a classical treatment is valid.

V. CONCLUSIONS

Exact solutions of equations of motion allow us to make statements that do not rely on approximations, and can give explicit insight into complex phenomena induced by, as considered here, radiation reaction. We have presented new exact solutions of the Landau-Lifshitz equation, for a longitudinally polarised electric field depending on a single lightlike coordinate, and of the ‘second-order-of-reduction’ Landau-Lifshitz equation (LL₂) for a plane wave.

We have used these solutions to examine the radiation-free direction hypothesis [30], that is the proposed universal approach of particle motion, in strong fields, to a direction which minimises radiation losses. We have explicitly confirmed this behaviour in our new solutions.

Many authors have found exact solutions to the LAD and/or LL equations in a number of field configurations, including fields depending only on time [28], constant electromagnetic fields [36, 52, 53], rotating electric fields [5, 53, 54], the Coulomb potential in the non-relativistic limit [55], and plane waves [45, 46]. These works, the majority of which predate the RFD hypothesis, contain implicit support for it. For example, in a rotating

electric field $\mathbf{E} = E_0 (\cos \omega t, \sin \omega t, 0)$, the particle momentum co-rotates with the electric force, lagging $e\mathbf{E}$ by an angle $\approx 90^\circ$ for small E_0 and $\sim E_0^{-3/4}$ for large E_0 [5, Sec. II.A]. Further, although the exact solution of LAD is not known even in a general constant field, its asymptotic behaviour is known [56], namely the particle worldline becomes confined to an eigen-2-plane of the field tensor; the same holds for the exact solution of LL_1 [36].

There are many examples of particle motion for which the Lorentz force law is integrable, or even superintegrable [57–59]. In future work it would be interesting to examine in detail how this integrability is affected by the addition of radiation-reaction terms in the LAD, LL and higher order, LL_n , forms. For example, the Lorentz force equation in a plane wave is superintegrable; in going to the Landau-Lifshitz equation (LL_1), one loses conserved quantities (i.e. π^+ is no longer conserved), which lowers the degree of integrability, while the second-order Landau-

Lifshitz equation (LL_2) becomes nonlinear and is only integrable in special cases.

Finding exact solutions of higher-reduction-of-order equations could shed light on the emergence of non-perturbative phenomena (runaways and acausal solutions), but this would require resummation [7]. This is particularly intriguing as resummation has recently been highlighted as being essential for fully understanding the behaviour of *quantum* dynamics in strong fields [27, 60, 61].

ACKNOWLEDGMENTS

The authors thank Ben King for useful comments and discussions. The authors are supported by the Leverhulme Trust (RE, AI, TH), grant RPG-2019-148, and the EPSRC (AI), grant EP/S010319/1.

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- [1] M. Abraham, *Theorie der Elektrizität* (Teubner, Leipzig, 1905).
- [2] H. A. Lorentz, *The Theory of Electrons* (Teubner, Leipzig, 1909).
- [3] P. A. M. Dirac, Proc. Roy. Soc. A **167**, 148 (1938).
- [4] A. Di Piazza, K. Z. Hatsagortsyan, and C. H. Keitel, Phys. Rev. Lett. **102**, 254802 (2009), arXiv:0810.1703 [physics.class-ph].
- [5] S. V. Bulanov, T. Z. Esirkepov, M. Kando, J. K. Koga, and S. S. Bulanov, Phys. Rev. E **84**, 056605 (2011).
- [6] S. Iso and S. Zhang, Phys. Rev. D **86**, 125019 (2012), arXiv:1207.7216 [hep-th].
- [7] S. Zhang, Prog. Theor. Exp. Phys **2013**, 123A01 (2013), arXiv:1303.7120 [hep-th].
- [8] M. Vranic, J. L. Martins, J. Vieira, R. A. Fonseca, and L. O. Silva, Phys. Rev. Lett. **113**, 134801 (2014), arXiv:1306.0766 [physics.plasm-ph].
- [9] V. Dinu, C. Harvey, A. Ilderton, M. Marklund, and G. Torgrimsson, Phys. Rev. Lett. **116**, 044801 (2016), arXiv:1512.04096 [hep-ph].
- [10] A. Gonoskov, S. Bastrakov, E. Efimenko, A. Ilderton, M. Marklund, I. Meyerov, A. Muraviev, A. Sergeev, I. Surmin, and E. Wallin, Phys. Rev. E **92**, 023305 (2015), arXiv:1412.6426 [physics.plasm-ph].
- [11] F. Li, V. K. Decyk, K. G. Miller, A. Tableman, F. S. Tsung, M. Vranic, R. A. Fonseca, and W. B. Mori, Accurately simulating nine-dimensional phase space of relativistic particles in strong fields (2020), arXiv:2007.07556 [physics.comp-ph].
- [12] J. M. Cole, K. T. Behm, E. Gerstmayr, T. G. Blackburn, J. C. Wood, C. D. Baird, M. J. Duff, C. Harvey, A. Ilderton, A. S. Joglekar, K. Krushelnick, S. Kuschel, M. Marklund, P. McKenna, C. D. Murphy, K. Poder, C. P. Ridgers, G. M. Samarin, G. Sarri, D. R. Symes, A. G. R. Thomas, J. Warwick, M. Zepf, Z. Najmudin, and S. P. D. Mangles, Phys. Rev. X **8**, 011020 (2018), arXiv:1707.06821 [physics.plasm-ph].
- [13] K. Poder, M. Tamburini, G. Sarri, A. Di Piazza, S. Kuschel, C. D. Baird, K. Behm, S. Bohlen, J. M. Cole, D. J. Corvan, M. Duff, E. Gerstmayr, C. H. Keitel, K. Krushelnick, S. P. D. Mangles, P. McKenna, C. D. Murphy, Z. Najmudin, C. P. Ridgers, G. M. Samarin, D. R. Symes, A. G. R. Thomas, J. Warwick, and M. Zepf, Phys. Rev. X **8**, 031004 (2018), arXiv:1709.01861 [physics.plasm-ph].
- [14] C. D. Baird, C. D. Murphy, T. G. Blackburn, A. Ilderton, S. P. D. Mangles, M. Marklund, and C. P. Ridgers, N. J. Phys. **21**, 053030 (2019), arXiv:1804.07725 [physics.plasm-ph].
- [15] C. F. Nielsen, J. B. Justesen, A. H. Sørensen, U. I. Uggerhøj, and R. Holtzapple, Phys. Rev. D **102**, 052004 (2020), arXiv:2008.10411 [hep-ex].
- [16] Extreme Light Infrastructure, <http://www.extreme-light-infrastructure.eu>.
- [17] C. N. Danson *et al.*, High Power Laser Sci. Eng. **7**, e54 (2019).
- [18] H. Abramowicz *et al.*, Conceptual Design Report for the LUXE Experiment (2021), arXiv:2102.02032 [hep-ex].
- [19] S. Meuren *et al.*, On Seminal HEDP Research Opportunities Enabled by Colocating Multi-Petawatt Laser with High-Density Electron Beams (2020), arXiv:2002.10051 [physics.plasm-ph].
- [20] D. A. Burton and A. Noble, Contemp. Phys. **55**, 110 (2014), arXiv:1409.7707 [physics.plasm-ph].
- [21] T. G. Blackburn, Rev. Mod. Plasma Phys. **4**, 5 (2020), arXiv:1910.13377 [physics.plasm-ph].
- [22] A. Gonoskov, A. Bashinov, I. Gonoskov, C. Harvey, A. Ilderton, A. Kim, M. Marklund, G. Mourou, and A. Sergeev, Phys. Rev. Lett. **113**, 014801 (2014), arXiv:1306.5734 [physics.plasm-ph].
- [23] S. V. Bulanov, T. Z. Esirkepov, J. K. Koga, S. S. Bulanov, Z. Gong, X.-Q. Yan, and M. Kando, J. Plasma Phys. **83**, 905830202 (2017).
- [24] G. Lehmann and K. H. Spatschek, Phys. Rev. E **84**, 046409 (2011).
- [25] C. Harvey, T. Heinzl, and M. Marklund, Phys. Rev. D **84**, 116005 (2011), arXiv:1110.0628 [physics.class-ph].
- [26] E. Gelfer, N. Elkina, and A. Fedotov, Sci. Rep. **8**, 6478 (2018).
- [27] T. Heinzl, A. Ilderton, and B. King, (2021),

- arXiv:2101.12111 [hep-ph].
- [28] G. N. Plass, *Rev. Mod. Phys.* **33**, 37 (1961).
- [29] L. D. Landau and E. M. Lifshitz, *The Classical Theory of Fields* (Butterworth-Heinemann, 1975).
- [30] A. Gonoskov and M. Marklund, *Phys. Plasmas* **25**, 093109 (2018).
- [31] T. N. Tomaras, N. C. Tsamis, and R. P. Woodard, *Phys. Rev. D* **62**, 125005 (2000), arXiv:hep-ph/0007166.
- [32] R. P. Woodard, *Nucl. Phys. B Proc. Suppl.* **108**, 165 (2002), arXiv:hep-th/0111282.
- [33] H. Wang, L. Shi, B. Lukyanchuk, C. Sheppard, and C. T. Chong, *Nature Photonics* **2**, 501 (2008).
- [34] P. A. M. Dirac, *Rev. Mod. Phys.* **21**, 392 (1949).
- [35] T. Fulton and F. Rohrlich, *Ann. Phys. (N.Y.)* **9**, 499 (1960).
- [36] Y. Yaremko, *J. Math. Phys.* **54**, 092901 (2013), arXiv:1412.1661 [hep-th].
- [37] D. Seipt and A. G. R. Thomas, *Plasma Phys. Control. Fusion* **61**, 074005 (2019), arXiv:1903.11463 [physics.plasm-ph].
- [38] J. Schwinger, L. L. DeRaad, K. Milton, and W.-Y. Tsai, *Classical electrodynamics* (Westview Press, 1998).
- [39] C. Eliezer, *Proc. Roy. Soc. A* **194**, 543 (1948).
- [40] G. Ford and R. O'Connell, *Phys. Lett. A* **157**, 217 (1991).
- [41] V. Krivitsky and V. Tsytovich, *Sov. Phys. Usp.* **34**, 250 (1991).
- [42] A. Higuchi, *Phys. Rev. D* **66**, 105004 (2002), [Erratum: *Phys. Rev. D* **69**, 129903 (2004)], arXiv:quant-ph/0208017.
- [43] A. Ilderton and G. Torgrimsson, *Phys. Rev. D* **88**, 025021 (2013), arXiv:1304.6842 [hep-th].
- [44] T. C. Mo and C. H. Papas, *Phys. Rev. D* **4**, 3566 (1971).
- [45] H. Heintzmann and M. Grewing, *Z. Phys.* **251**, 77 (1972).
- [46] A. Di Piazza, *Lett. Math. Phys.* **83**, 305 (2008).
- [47] A. Polyanin and V. Zaitsev, *Handbook of Exact Solutions for Ordinary Differential Equations*, 2nd ed. (Chapman and Hall/CRC, 2002).
- [48] T. W. B. Kibble, *Phys. Rev.* **138**, B740 (1965).
- [49] T. C. Adorno, S. P. Gavrilov, and D. M. Gitman, *Eur. Phys. J. C* **76**, 447 (2016), arXiv:1605.09072 [hep-th].
- [50] A. Di Piazza, *Phys. Lett. B* **782**, 559 (2018), arXiv:1804.01160 [physics.plasm-ph].
- [51] R. M. Corless, G. H. Gonnet, D. E. G. Hare, D. J. Jeffrey, and D. E. Knuth, *Adv. Comput. Math.* **5**, 329 (1996).
- [52] G. Ares de Parga and R. Mares, *J. Math. Phys.* **40**, 4807 (1999).
- [53] T. Nakamura, *Phys. Rev. E* **102**, 033210 (2020).
- [54] R. Rivera and D. Villarroel, *Phys. Rev. E* **66**, 046618 (2002).
- [55] S. G. Rajeev, *Ann. Phys. (N.Y.)* **323**, 2654 (2008), arXiv:0801.0238 [hep-th].
- [56] P. O. Kazinski and M. A. Shipulya, *Phys. Rev. E* **83**, 066606 (2011), arXiv:1012.5728 [hep-th].
- [57] W. Miller, Jr., S. Post, and P. Winternitz, *J. Phys. A* **46**, 423001 (2013), arXiv:1309.2694 [math-ph].
- [58] V. G. Bagrov and D. Gitman, *The Dirac Equation and its Solutions*, De Gruyter Studies in Mathematical Physics, Vol. 4 (De Gruyter, 2014).
- [59] T. Heinzl and A. Ilderton, *Phys. Rev. Lett.* **118**, 113202 (2017), arXiv:1701.09166 [hep-ph].
- [60] A. A. Mironov, S. Meuren, and A. M. Fedotov, *Phys. Rev. D* **102**, 053005 (2020), arXiv:2003.06909 [hep-th].
- [61] J. P. Edwards and A. Ilderton, *Phys. Rev. D* **103**, 016004 (2021), arXiv:2010.02085 [hep-ph].